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On the stability of density stratified flow below a ponded surface: comprehensive version

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Abstract

*Flooding of coastal areas with seawater often leads to density stratification. The stability of the density-depth profile in a porous medium initially saturated with a fluid of density ρ_f after flooding with a salt solution of higher density ρ_s is analyzed. The standard convection/diffusion equation subject to the so-called Boussinesq approximation is used. The depth of the porous medium is assumed to be infinite in the analytical approaches and finite in the numerical simulations. Two cases are distinguished: the laterally unbounded **CASE A** and the laterally bounded **CASE B**. The ratio of the diffusivity and the density difference ($\rho_s - \rho_f$) induced gravitational shear flow is an intrinsic length scale of the problem. In the unbounded **CASE A**, this geometric length scale is the only length scale and using it to write the problem in dimensionless form results in a formulation with Rayleigh number $R=1$. In the bounded **CASE B**, the lateral geometry provides another length scale. Using this geometrical length scale to write the problem in dimensionless form results in a formulation with a Rayleigh number R given by the ratio of the geometric and intrinsic length scales.*

*For both **CASE A** and **CASE B**, the well-known Boltzmann similarity solution provides the ground state. Three analytical approaches are used to study the stability of this ground state, the first two based on the linearized perturbation equation for the concentration and the third based on the full nonlinear equation. For the first linear approach, the surface spatial density gradient is used as an approximation of the entire background density profile. This results in a crude estimate of the L^2 -norm of the concentration showing that the perturbation at first grows, but eventually decays in time.*

*For the other two approaches, the full ground state solution is used, although for the second linear approach subject to the restriction that the ground state slowly evolves in time (the so-called 'frozen profile approximation'). Just like the ground state, the resulting eigenvalue problems can be written in terms of the Boltzmann variable. The linearized stability approach holds only for infinitesimal small perturbations, whereas the nonlinear, variational energy approach is not subject to such a restriction. The results for all three approaches can be expressed in terms of Boltzmann \sqrt{t} transformed relationships between the system Rayleigh number and perturbation wave number. The results of the linear and nonlinear approaches are surprisingly close to each other. Based on the system Rayleigh number, this allows delineation of systems that are unconditionally stable, marginally stable, or episodically unstable. These analytical predictions are confirmed by direct two-dimensional numerical simulations, which also show the details of the episodic instabilities as function of the wave number for case **CASE A** and the wave number and Rayleigh number for **CASE B**. A numerical example of the effect of a layer with*

low permeability is also shown. Using typical values of the physical parameters, the analytical and numerical results are interpreted in terms of dimensional length and time scales. In particular an explicit stability criterion is given for vertical column experiments.

1. INTRODUCTION

Stability of density stratified fluids has been studied for more than 100 years. In the book 'The self-made tapestry', Ball (2001) gives in chapter 7 a lucid description of the history of the analysis of pattern formation in such fluids. The extension to fluids in porous media is nearly as old and has been reviewed thoroughly by Straughan (2008). At first, stability was explored primarily by analytical methods applied to the linearized perturbation equations. Later it became feasible to use analytical and numerical methods to study the full nonlinear equations.

Employing those modern methods, van Duijn et al. (2002) analyzed gravitational stability of a saline boundary layer below an evaporating salt lake. They compared stability criteria obtained by three different methods: the method of linearized stability; the traditional energy method using as constraint the integrated Darcy equation; an alternative energy method using as constraint the point-wise Darcy equation. The two constraints are, respectively, referred to as integral constraint and differential constraint. They showed that the integral constraint gives a stability bound of the Rayleigh number equal to the square of the first root of the Bessel function J_0 , in agreement with previous numerical results of Homsy and Sherwood (1976). The differential constraint gave results that were in excellent agreement with experiments by Wooding et al. (1997a,b). The stability of the saline boundary layer formed by upward flow was treated in greater detail by Pieters (2004) and Pieters and Schuttelaars (2008).

In this paper we use the same approach to analyze stability following flooding with salt water of a soil initially saturated with less saline water. In the natural environment, such flooding is quite common. Marine transgression is a dominant feature of Holocene geo-hydrological history, starting almost 12 millennia ago. For example, see Raats (2015) for a historical perspective, the Dutch coastal region is shaped by marine transgression and human interference in response to this. In the region where now the fresh-water IJsselmeer and the IJsselmeerpolders are located, there was in Roman times the fresh-water Flevomeer, later known as Almere. Due to marine transgression, the inland lake gradually evolved into the inland Zuiderzee, which stood in open connection with the North Sea. From about 1600 AD onward, the bottom of Zuiderzee became more and more saline. With the double purpose of protection from floods and the creation of new polders, in 1932 the tidal inlet to the Zuiderzee was closed off by a dam. In a few years the brackish Zuiderzee changed into the fresh-water IJsselmeer and as a result of that the bottom started to desalinize. That process is still continuing.

In the 1930s and early 1940s, again following Raats (2015), a group of civil engineers inspired by the mathematical physicist J. M. Burgers did pioneering work on transport of salts across the water-sediment interface. They formulated the linear convection-dispersion equation and solved this equation for appropriate boundary conditions. The calculated salinity profiles in the top 10 meters were mostly in good agreement with profiles measured in the 1950s by W.H. van der Molen (see Raats (2015) and the references therein). The spatial distribution of the measured salinity profiles over a large area, such as the North-East polder, were used as an indicator of the spatial distribution of the flow velocities. However, not all field measurements were in line with the computations. In a few scattered places in the North-East polder higher salinities were found at depths of 10–15 m, far deeper than predicted and measured in other places. In some cases this could reflect seepage of water from the former Zuiderzee to lower lying polders along the coast. In others there was a highly permeable Pleistocene deposit reaching the surface. Van der Molen speculated that for the latter 'this phenomenon is probably due to convective currents in the bottom of the Zuiderzee between 1600 and 1931 A.D.'. Specifically, he noted that the small difference in density between the fresh water present in the soil and the supernatant seawater is

sufficient to cause convection currents. This is in line with the theory and computations presented in this paper.

On the time scale of centuries and under certain circumstances, marine transgression may cause rapid salinization of entire aquifers. In northwest Germany, Holocene transgressions of a few thousands of years have brought salt water of corresponding age to a depth of over 200 m (Mull and Battermann (1980)). Earlier Geirnaert (1973) had reported similar observations along the Dutch coast. In view of such results, it may seem surprising that at many other places all around the world fresh and brackish waters have been found beneath saline waters on the continental shelves (Post et al., 2013). Wherever this occurs, the deep-lying fresh water is invariably protected from invading ocean waters by sediments with a low hydraulic conductivity. Post and Simmons (2009) illustrate by means of sand tank laboratory experiments and numerical modeling how low-permeability lenses protect fresh water from mixing with overlying saline water. This is also in line with the theory presented in this paper.

On a seasonal time scale, exchange of substances between fresh water in sediments and supernatant saline water may occur. Smetacek et al. (1976) reported that in the water of the Kiel Bight high nutrient concentrations and low oxygen concentrations were found following an influx of higher density water. They suggested that this could be related to density induced natural convection in the sediments. Webster et al. (1996) used a computational model and laboratory experiments to demonstrate that gravitational convection can make an important contribution to the exchange of water and solutes between sediments and a supernatant water column in regions subject to significant temporal variations in salinity, such as estuaries. In effect, they verified the suggestion put forward by Smetacek et al. (1976).

In coastal regions, tsunamis often cause flooding by seawater and natural convection may then strongly contribute to the mixing of this seawater with the underlying fresh water. Illangasekare et al. (2006) described the impact of the 2004 tsunami in Sri Lanka on groundwater resources. They included results of an experiment in a $53 \times 30 \times 2.7$ cm PlexiglasTM tank filled with glass beads. Vithanage et al. (2008) used a laboratory model to study natural convection in coastal aquifers following flooding by sea water.

The analysis of the stability of density stratified flow below a ponded surface in this paper has important connections with a pioneering paper by Wooding (1962b). In that paper, Wooding uses the method of linearized stability to analyze the behavior of an initially sharp, horizontal interface separating two miscible fluids in an infinitely long, vertical, porous column. In the upper half of the column the fluid has a higher density than in the lower half. He studied the time dependence of the linearized equations by means of an expansion in terms of Hermite polynomials. In spite of the rather primitive numerical tools available in those days, the approach undertaken by Wooding in this and other papers - both theoretical and experimental: see e.g. Wooding (1959, 1960, 1962a, 1969); Wooding et al. (1997a,b) - is still a source of inspiration for those studying stability problems in porous media. For this reason we dedicate this paper in his honor.

Long before we all were made energy conscious in the 1970s, already in the mid 1950s Robin Wooding (1926–2007) was concerned with geothermal problems. He started his research in this area at the DSIR Applied Mathematics Laboratory at Wellington, New Zealand in the mid 1950s, continued it with a DSIR National Research Fellowship at Emmanuel College of the University of Cambridge. There he obtained in 1960 his PhD-degree under the supervision of Dr. P.G. Saffman in the group of Professor G.I. Taylor. He returned to DSIR, from where he moved in 1963 on to CSIRO at Canberra to work on overland flow and near-surface, atmospheric turbulence. On a Senior Research Fellowship, in 1968 he visited the California Institute of Technology and there he wrote Wooding (1969). In 1970/71 he took leave without pay to spend a year at The Johns Hopkins University and the University of Wisconsin to work on flow and transport in porous media, after which he decided to return not to CSIRO but to DSIR, where he stayed until he retired in 1987. He then returned again to Canberra as Honorary Fellow at the CSIRO Pye Laboratory. There for the last two decades of his life his main interest was the salt lake problem mentioned above.

After this introduction, the outline of this paper is as follows. The mathematical model for studying the stability of density stratified flow below a ponded surface is formulated in Section 2, where we distinguish between laterally unconfined and confined domains. A special solution of this model is presented in Section 3. This solution describes the formation of a one-dimensional salt layer that penetrates into the subsurface by diffusion/dispersion. It is the main goal of this paper to investigate the gravitational (in)stability of this layer. In particular, we study the influence of the system's dimensionless Rayleigh number and the role of time. To this end we perturb the special solution and derive the corresponding perturbation equations. This is done in Section 4. Next, in Section 5, we investigate the stability by two different methods, namely the standard method of linearized stability, where we employ two approaches, and the energy method with differential constraint. Stability results for homogeneous soils are discussed in Section 6.1 and for non-homogeneous (layered) soils in Section 6.2. An interpretation of the results in terms of the physical parameters of the problem is given in Section 7, and conclusions are presented in Section 8.

2. PROBLEM FORMULATION

We consider an isotropic porous medium occupying either

CASE A: the laterally unbounded half-space $\Omega = \Omega_A = \{z > 0\}$, or

CASE B: the three dimensional, laterally confined, semi-infinite region $\Omega = \Omega_B = \{(x, y, z) : -L_{x,y} < x, y < L_{x,y}, z > 0\}$.

Here z is the vertical coordinate pointing downwards. We assume that the permeability κ is constant in the x, y direction, i.e. $\kappa = \kappa(z)$ only, and that the porosity ϕ is constant throughout Ω .

Adopting the Boussinesq approximation (Oberbeck (1879), Boussinesq (1903), Chandrasekhar (1981), Landman and Schotting (2007)), we consider the equations

$$\phi \frac{\partial \rho}{\partial t} + \nabla \cdot (\mathbf{U} \rho) - D \Delta \rho = 0, \quad (1a)$$

$$-\frac{\kappa(z)}{\mu} \nabla (\tilde{P} - \rho_f g z) + \frac{\kappa(z)}{\mu} (\rho - \rho_f) g \mathbf{e}_z = \mathbf{U}, \quad (1b)$$

$$\nabla \cdot \mathbf{U} = 0, \quad (1c)$$

in Ω and for $t > 0$. In these equations, ρ denotes the fluid density, \mathbf{U} the Darcy velocity, and \tilde{P} the fluid pressure. These are the unknowns in the problem that need to be determined. The (given) physical parameters are: D (dispersion/diffusion coefficient), μ (fluid viscosity), g (gravity constant), and ρ_f (fresh water density). Both D and μ are assumed constant throughout this paper.

The boundary conditions at the top of the porous medium are given by:

$$\rho = \rho_s \quad \text{and} \quad \tilde{P} = \rho_s g h \quad \text{at } z = 0, \text{ for } t > 0, \quad (2a)$$

where ρ_s is the salt (sea) water density and h the height of the overlying salt water layer. Boundary condition (2a) is a simplification of what happens in practice. The laboratory experiments by Webster et al. (1996) show that growing instabilities induce local outflow from the porous medium into the ponded layer. It is assumed in this paper that outflow across the top of the porous medium $\{z = 0\}$ has no effect on the density of the fluid in the overlying layer (which is ρ_s for all $t > 0$).

The initial condition is given by:

$$\rho = \rho_i \quad \text{in } \Omega, \text{ at } t = 0, \quad (2b)$$

where the initial density satisfies $\rho_f \leq \rho_i \leq \rho_s$.

Remark 1 (i) The results for **CASE B** can be generalized to arbitrarily shaped bounded cross-sections. In Section 7 we consider the case of a circular cross-section as well.

(ii) Since Darcy's law contains the gradient of the pressure, one may replace \tilde{P} by $\tilde{P} + \tilde{P}_0$, with \tilde{P}_0 constant in space, and obtain the same discharge and salt distribution. Hence the specific value of \tilde{P} at the top $z = 0$ does not play a role. In other words, the discharge and salt distribution do not depend on the height h of the ponded layer. In essence this is due to the assumed constraint of incompressibility of the fluid.

It is assumed that $\rho_i \rightarrow \rho_f$ as $z \rightarrow \infty$, implying that $\rho \rightarrow \rho_f$ as $z \rightarrow \infty$ for all $t > 0$. Thus ρ_f acts as a far field reference value. We further assume that as $z \rightarrow \infty$, the pressure gradient force is balanced by the gravitational force. Thus, we consider those solutions of (1) that satisfy:

$$\rho \rightarrow \rho_f \text{ and } \frac{\partial \tilde{P}}{\partial z} \rightarrow \rho_f g \quad \text{as } z \rightarrow \infty, \text{ for all } t > 0. \quad (2c)$$

For the horizontally bounded **CASE B**, it is assumed that on lateral boundary Γ_v (see Figure 2) the fluxes of water and solute vanish:

$$\mathbf{U} \cdot \mathbf{n} = 0, \quad (\mathbf{U}\rho - D\nabla\rho) \cdot \mathbf{n} = 0 \quad \text{along } \Gamma_v, \text{ for } t > 0, \quad (2d)$$

where \mathbf{n} is the unit vector on Γ_v .

Using the constant densities ρ_s and ρ_f we define:

$$C = \frac{\rho - \rho_f}{\rho_s - \rho_f} \quad \text{and} \quad P = \tilde{P} - \rho_f g z. \quad (3)$$

Here P is the pressure based on the far field reference density ρ_f . Introducing (3) into (1) and (2) gives

$$\phi \frac{\partial C}{\partial t} + \nabla \cdot (\mathbf{U}C - D\nabla C) = 0, \quad (4a)$$

$$-\frac{\kappa(z)}{\mu} \nabla P + \frac{\kappa(z)}{\mu} (\rho_s - \rho_f) g \mathbf{e}_z = \mathbf{U}, \quad (4b)$$

$$\nabla \cdot \mathbf{U} = 0, \quad (4c)$$

$$C = C_i \quad \text{in } \Omega, \text{ at } t = 0, \quad (5a)$$

$$C = 1, \quad P = \rho_s g h \quad \text{at } z = 0, \text{ for } t > 0, \quad (5b)$$

$$C \rightarrow 0, \quad \frac{\partial P}{\partial z} \rightarrow 0 \quad \text{as } z \rightarrow \infty, \text{ for } t > 0, \quad (5c)$$

and for **CASE B** also

$$\mathbf{U} \cdot \mathbf{n} = 0, \quad (\mathbf{U}C - D\nabla C) \cdot \mathbf{n} = 0 \quad \text{along } \Gamma_v, \text{ for } t > 0. \quad (5d)$$

Next we introduce the characteristic scales for length (L_c), time (T_c), pressure (P_c), discharge (Q_c) and permeability ($\kappa_c = \kappa(0)$). Using these scales, we substitute the dimensionless variables

$$(x^*, y^*, z^*) = \frac{(x, y, z)}{L_c}, \quad t^* = \frac{t}{T_c}, \quad P^* = \frac{P}{P_c}, \quad \mathbf{U}^* = \frac{\mathbf{U}}{Q_c}, \quad \kappa^* = \frac{\kappa}{\kappa_c}, \quad (6)$$

into the equations (4) and conditions (5) to obtain in the scaled domain Ω^* :

$$\frac{\phi L_c^2}{T_c D} \frac{\partial C}{\partial t^*} + R \nabla^* \cdot (\mathbf{U}^* C) - \Delta^* C = 0, \quad (7a)$$

$$-\frac{\kappa_c \kappa^*(z^*) P_c}{\mu L_c} \nabla^* P^* + \frac{\kappa_c \kappa^*(z^*)}{\mu} (\rho_s - \rho_f) g C \mathbf{e}_z = Q_c \mathbf{U}^*, \quad (7b)$$

$$\nabla \cdot \mathbf{U}^* = 0, \quad (7c)$$

subject to

$$C = C_i \quad \text{in } \Omega^*, \text{ at } t^* = 0, \quad (8a)$$

$$C = 1, \quad P^* = \frac{\rho_s g h}{P_c} \quad \text{at } z^* = 0, \text{ for } t^* > 0, \quad (8b)$$

$$C \rightarrow 0, \quad \frac{\partial P^*}{\partial z^*} \rightarrow 0 \quad \text{as } z^* \rightarrow \infty, \text{ for } t^* > 0, \quad (8c)$$

and for **CASE B** also

$$\mathbf{U}^* \cdot \mathbf{n} = 0, \quad (R \mathbf{U}^* C - \nabla^* C) \cdot \mathbf{n} = 0 \quad \text{along } \Gamma_v^*, \text{ for } t > 0. \quad (8d)$$

The transformed region Ω^* is given by

$$(\text{CASE A}) \quad \Omega_A^* = \{z^* > 0\}, \quad (9a)$$

$$(\text{CASE B}) \quad \Omega_B^* = \{(x^*, y^*, z^*) : |x^*| < \frac{L_x}{L_c}, |y^*| < \frac{L_y}{L_c}, z^* > 0\}, \quad (9b)$$

and in (7) and (8) the Rayleigh number R is given by

$$R = \frac{Q_c L_c}{D}. \quad (10)$$

Note that in (7b) we introduced

$$\kappa^*(z^*) = \frac{\kappa(L_c z^*)}{\kappa_c}. \quad (11)$$

Balancing the first and third term in transport equation (7a) gives for the time scale T_c :

$$T_c = \frac{\phi L_c^2}{D}, \quad (12a)$$

and, similarly, balancing all terms in Darcy's law (7a) gives the pressure scale P_c :

$$P_c = \frac{\mu L_c Q_c}{\kappa_c}, \quad (12b)$$

and discharge scale Q_c :

$$Q_c = \frac{\kappa_c (\rho_s - \rho_f) g}{\mu}. \quad (12c)$$

The discharge scale Q_c can be interpreted as a density difference induced, gravitational shear flow (de Josselin de Jong, 1960).

The time scale T_c and pressure scale P_c are known once the length scale L_c has been selected. In the laterally unbounded **CASE A**, there is no obvious length scale in terms of the spatial dimensions of the problem. Therefore we use the intrinsic length scale $L_c = D/Q_c$. In the laterally bounded

Table 1: Overview of the definitions of the **CASE A** and **CASE B** scales.

CASE A	CASE B
$Q_c = \frac{\kappa_c(\rho_s - \rho_f)g}{\mu}$	
$L_c = \frac{D}{Q_c}$	$L_c = \frac{L_x}{\pi}$
$T_c = \frac{\phi D}{Q_c^2}$	$T_c = \frac{\phi L_c^2}{D} = \frac{\phi L_x^2}{\pi^2 D}$
$P_c = \frac{\mu D}{\kappa_c}$	$P_c = \frac{\mu L_c Q_c}{\kappa_c} = \frac{\mu L_x Q_c}{\pi \kappa_c}$
$R = 1$	$R = \frac{L_c}{D/Q_c} = \frac{L_x}{\pi D/Q_c}$

CASE B, we choose $L_c = L_x/\pi$ or some other characteristic horizontal length scale. The factor π is for notational and technical convenience. All scales are summarized in Table 1.

For **CASE B**, the Rayleigh number is, apart from the factor π , the ratio of the characteristic geometric length scale L_x and the intrinsic length scale D/Q_c . Its value may vary considerably due to L_x and, in particular, κ_c (see Section 7).

The scaled region Ω^* is given by

$$(\text{CASE A}) \quad \Omega_A^* = \{z^* > 0\}, \quad (13a)$$

$$(\text{CASE B}) \quad \Omega_B^* = \{(x^*, y^*, z^*) : |x^*| < \pi, |y^*| < \frac{L_y}{L_x} \pi, z^* > 0\}. \quad (13b)$$

Thus after scaling we obtain the dimensionless form of the equations, dropping the superscript $*$,

$$\frac{\partial C}{\partial t} + R \nabla \cdot (\mathbf{U} C) - \Delta C = 0, \quad (14a)$$

$$\frac{1}{\kappa(z)} \mathbf{U} = -\nabla P + C \mathbf{e}_z, \quad (14b)$$

$$\nabla \cdot \mathbf{U} = 0, \quad (14c)$$

in Ω and for $t > 0$, subject to

$$C = C_i \quad \text{in } \Omega, \text{ at } t = 0, \quad (15a)$$

$$C = 1, \quad P = \frac{\rho_s g h}{P_c} \quad \text{at } z = 0, \text{ for } t > 0, \quad (15b)$$

$$C \rightarrow 0, \quad \frac{\partial P}{\partial z} \rightarrow 0 \quad \text{as } z \rightarrow \infty, \text{ for } t > 0, \quad (15c)$$

and for **CASE B** also

$$\mathbf{U} \cdot \mathbf{n} = 0, \quad (R \mathbf{U} C - \nabla C) \cdot \mathbf{n} = 0 \quad \text{along } \Gamma_v, \text{ for } t > 0. \quad (15d)$$

The transformed region Ω is given by

$$(\text{CASE A}) \quad \Omega_A = \{z > 0\}, \quad (16a)$$

$$(\text{CASE B}) \quad \Omega_B = \{(x, y, z) : |x| < \pi, |y| < \frac{L_y}{L_x} \pi, z > 0\}, \quad (16b)$$

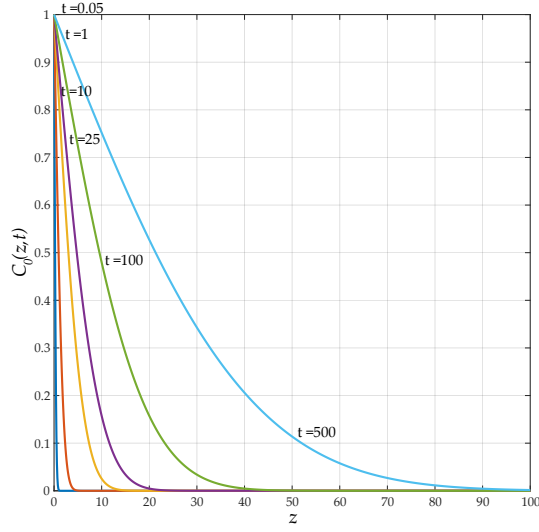


Fig. 1: Ground state profiles for various values of time.

Remark 2 Suppose one compares different values of R in **CASE B**. These differences are due to differences in κ_c and L_x . Note that if L_x changes, then $\kappa(z)$ changes accordingly, because of the scaling of z . This is expressed by (11). For **CASE A**, the κ_c only appears in the length scale.

3. THE GROUND STATE SOLUTION

In this paper we investigate the stability of the particular solution of (14) subject to (15) for $C_i \equiv 0$, i.e. we assume that the initial density ρ_i is equal to the far field density ρ_f . We call this particular solution the ground state and denote it by $\{C_0, U_0, P_0\}$. Because the boundary and initial conditions do not depend on the horizontal coordinates x and y , the ground state depends only on z and t . Thus $\{C_0, U_0, P_0\}$ is the solution of

$$\frac{\partial C_0}{\partial t} + R \frac{\partial U_0 C_0}{\partial z} - \frac{\partial^2 C_0}{\partial z^2} = 0, \quad (17a)$$

$$\frac{1}{\kappa(z)} U_0 = -\frac{\partial P_0}{\partial z} + C_0, \quad (17b)$$

$$\frac{\partial U_0}{\partial z} = 0, \quad (17c)$$

subject to

$$C_0 = 1 \text{ and } P_0 = \frac{\rho_s g h}{P_c} \quad \text{at } z=0, \text{ for } t>0, \quad (18a)$$

$$C_0 \rightarrow 0 \text{ and } \frac{\partial P_0}{\partial z} \rightarrow 0 \quad \text{as } z \rightarrow \infty, \text{ for } t>0, \quad (18b)$$

$$C_0 = 0 \quad \text{in } 0 < z < \infty \text{ and for } t=0. \quad (18c)$$

Applying (18b) in (17b) shows that $U_0 = 0$ for $z \rightarrow \infty$. In view of (17c) this means that $U_0 = 0$ for all z . Using this in (17) gives:

$$\frac{\partial C_0}{\partial t} = \frac{\partial^2 C_0}{\partial z^2}, \quad (19a)$$

$$\frac{\partial P_0}{\partial z} = C_0. \quad (19b)$$

Solving (19a) subject to (18) yields

$$C_0(z, t) = \operatorname{erfc}\left(\frac{z}{2\sqrt{t}}\right). \quad (20a)$$

This one-dimensional solution describes the formation of a salt layer that enters the subsurface by diffusion. Integrating (19b) subject to (18a) and (20a) results in

$$P_0(z, t) = \frac{\rho_s g h}{P_c} + 2\sqrt{t} \int_0^{\frac{z}{2\sqrt{t}}} \operatorname{erfc}(\zeta) d\zeta. \quad (20b)$$

Typical profiles for various $t > 0$ are shown in Figure 1. Note that

$$\lim_{t \rightarrow \infty} P_0(z, t) = \frac{\rho_s g h}{P_c} + z \quad \text{for each fixed } z > 0,$$

or in terms of the original variables

$$\lim_{t \rightarrow \infty} \tilde{P}(z, t) = \rho_s g(z + h) \quad \text{for each fixed } z > 0. \quad (21)$$

Equation (21) reflects the fact that ponding with a layer of fluid having constant density ρ_s , results in the halfspace Ω being filled with stagnant fluid of density ρ_s as $t \rightarrow \infty$.

4. THE PERTURBATION EQUATIONS

To investigate the stability of the ground state $\{C_0, \mathbf{U}_0, P_0\}$ we write the original variables as

$$C = C_0 + c, \quad \mathbf{U} = \mathbf{U}_0 + \mathbf{u}, \quad P = P_0 + p. \quad (22)$$

where $\mathbf{u} = (u, v, w)^T$. Here c , \mathbf{u} and p are the perturbations whose behaviour we need to study. For this we have to first write the perturbation equations. Substituting (22) in equations (14), the initial condition (15a) and the boundary conditions (15b) gives

$$\frac{\partial(C_0 + c)}{\partial t} = -\nabla \cdot ((C_0 + c)(\mathbf{U}_0 + \mathbf{u})) + \Delta(C_0 + c), \quad (23a)$$

$$\frac{1}{\kappa(z)}(\mathbf{U}_0 + \mathbf{u}) = -\nabla(P_0 + p) + (C_0 + c)\mathbf{e}_z, \quad (23b)$$

$$\nabla \cdot (\mathbf{U}_0 + \mathbf{u}) = 0, \quad (23c)$$

subject to

$$C_0 + c = C_i \quad \text{in } \Omega, \text{ for } t = 0, \quad (24a)$$

$$C_0 + c = 1, \quad P_0 + p = 1 \quad \text{at } z = 0, \text{ for } t > 0, \quad (24b)$$

$$C_0 + c \rightarrow 0, \quad \frac{\partial(P_0 + p)}{\partial z} \rightarrow 0 \quad \text{as } z \rightarrow \infty, \text{ for } t > 0, \quad (24c)$$

and for **CASE B** also

$$(\mathbf{U}_0 + \mathbf{u}) \cdot \mathbf{n} = 0, \quad (R(\mathbf{U}_0 + \mathbf{u})(C_0 + c) - \nabla(C_0 + c)) \cdot \mathbf{n} = 0 \quad \text{along } \Gamma_v, \text{ for } t > 0. \quad (24d)$$

In view of the properties of the ground state solution, there results for the perturbations the equations

$$\frac{\partial c}{\partial t} + R w \frac{\partial C_0}{\partial z} - \Delta c + R \mathbf{u} \cdot \nabla C_0 = 0, \quad (25a)$$

$$\frac{1}{\kappa(z)} \mathbf{u} + \nabla p - c \mathbf{e}_z = \mathbf{0}, \quad (25b)$$

$$\nabla \cdot \mathbf{u} = 0, \quad (25c)$$

and the boundary conditions

$$c = 0, \quad p = 0 \quad \text{at } z=0, \text{ for } t>0, \quad (26a)$$

$$c \rightarrow 0, \quad \frac{\partial p}{\partial z} \rightarrow 0 \quad \text{as } z \rightarrow \infty \text{ and } t>0, \quad (26b)$$

and for **CASE B** also

$$\mathbf{u} \cdot \mathbf{n} = 0, \quad \nabla c \cdot \mathbf{n} = 0 \quad \text{along } \Gamma_v, \text{ for } t>0. \quad (26c)$$

Multiplying (25b) by $\kappa(z)$ and applying the divergence gives

$$\nabla \cdot \mathbf{u} = 0 = \nabla \cdot (-\kappa(z)\nabla p + \kappa(z)c\mathbf{e}_z).$$

Hence

$$\frac{\partial}{\partial z} \left(-\kappa(z) \frac{\partial p}{\partial z} + \kappa(z)c \right) - \kappa(z)\Delta_{\perp} p = 0,$$

where $\Delta_{\perp} = \frac{\partial^2}{\partial x^2} + \frac{\partial^2}{\partial y^2}$ denotes the horizontal Laplacian. Substituting the vertical component w of the Darcy equation gives

$$\frac{\partial w}{\partial z} = \kappa(z)\Delta_{\perp} p. \quad (27)$$

Using this identity, the pressure boundary condition in (26a) implies for w the Neumann condition

$$\frac{\partial w}{\partial z} = 0 \quad \text{at } z=0, \text{ for } t>0. \quad (28)$$

Following Lapwood (1948), we take the Laplacian of the vertical component of the perturbation Darcy equation (25b). This gives

$$\Delta \left(\frac{1}{\kappa(z)} w \right) = -\frac{\partial}{\partial z} \Delta p + \Delta c. \quad (29)$$

Further, taking the divergence of the perturbed Darcy equation (25b) and using (25c),

$$\nabla \cdot \left(\frac{1}{\kappa(z)} \mathbf{u} \right) = \mathbf{u} \cdot \nabla \frac{1}{\kappa(z)} = w \frac{d}{dz} \frac{1}{\kappa(z)} = -\Delta p + \frac{\partial c}{\partial z},$$

yielding upon differentiation with respect to z

$$\frac{\partial}{\partial z} \left(w \frac{d}{dz} \frac{1}{\kappa(z)} \right) = -\frac{\partial}{\partial z} \Delta p + \frac{\partial^2 c}{\partial z^2}. \quad (30)$$

Subtracting (30) from (29) results in

$$\Delta \left(\frac{1}{\kappa(z)} w \right) - \frac{\partial}{\partial z} \left(w \frac{d}{dz} \frac{1}{\kappa(z)} \right) = \Delta_{\perp} c,$$

or

$$\nabla \cdot \left(\frac{1}{\kappa(z)} \nabla w \right) = \Delta_{\perp} c. \quad (31)$$

The Neumann condition (28) and identity (31) will be used frequently in the next sections.

5. STABILITY ANALYSES

In the unbounded **CASE A**, we consider x, y -periodic solutions of the perturbation equations with respect to x, y -periodic cells of the form

$$\Omega_A := \{(x, y, z) : (x, y) \in \mathcal{T}, z > 0\} , \quad (32)$$

where $\mathcal{T} \subset \mathbb{R}^2$ is an element (tile) of the horizontal periodic structure. This implies a redefinition of Ω_A as introduced in Section 2. As noted in Chandrasekhar (1981) or Straughan (2008), \mathcal{T} can be a rectangle (or square), a strip or a hexagon. In this paper the focus is on nonlinear stability and on the onset of infinitesimal small perturbations (linear approach). Details about the characterization of the periodic structure as for instance in Mielke (2002), who studied the spatial structure in the Rayleigh–Bénard problem, are beyond the scope of this paper. Some aspects, though, are discussed in Appendix A.

In the linear approach, the horizontal periodicity is described by a function $f = f(x, y)$ that satisfies

$$\Delta_{\perp} f + a^2 f = 0 \quad \text{in } \mathbb{R}^2 , \quad (33)$$

where a is a characteristic wavenumber. If \mathcal{T} is a hexagon with side L one has (see Chandrasekhar (1981))

$$f(x, y) = \cos \frac{4\pi}{3L} \left(\frac{\sqrt{3}}{2}x + \frac{1}{2}y \right) + \cos \frac{4\pi}{3L} \left(\frac{\sqrt{3}}{2}x - \frac{1}{2}y \right) + \cos \frac{4\pi}{3L} y , \quad (34a)$$

and

$$a = \frac{4\pi}{3L} . \quad (34b)$$

In this paper we visualize the horizontal periodicity in the form of rectangles $\mathcal{T} = (0, \ell_x) \times (0, \ell_y)$, where ℓ_x and ℓ_y are for the moment unspecified. This yields periodic cells of the form

$$\Omega_A := \{(x, y, z) : |x| < \ell_x, |y| < \ell_y, z > 0\} . \quad (35)$$

In (35), ℓ_x, ℓ_y are dimensionless lengths. A sketch of a two-dimensional cross-section of the flow domain in the x, z -plane is shown in Figure 2.

In the x, y -periodic assumption, periodic boundary conditions are applied along the vertical boundary of Ω_A , which we denote by Γ_v . This means that c , \mathbf{u} , and p can be extended in a periodic way over the whole Ω , and belong to $C^\infty(\Omega)$.

In the bounded **CASE B** we consider the perturbation equations in the flow domain $\Omega = \Omega_B$, which we write as

$$\Omega_B = \{(x, y, z) : |x| < \pi, |y| < \ell\pi, z > 0\} , \quad (36)$$

where $\ell = \frac{L_y}{L_x}$ is a dimensionless length ratio. Here we use no-flow and no-transport conditions along the vertical boundary Γ_v .

For computational purposes, see Section 6.1.3, we use the depth-truncated versions of Ω_A and Ω_B :

$$\Omega_A^d := \{(x, y, z) : |x| < \ell_x, |y| < \ell_y, 0 < z < d\} , \quad (37)$$

and

$$\Omega_B^d := \{(x, y, z) : |x| < \pi, |y| < \ell\pi, 0 < z < d\} . \quad (38)$$

where d denotes the dimensionless depth.

In Subsection 5.1 we consider the method of linearized stability, valid for small perturbations, and in Subsection 5.2 the nonlinear, variational energy method.

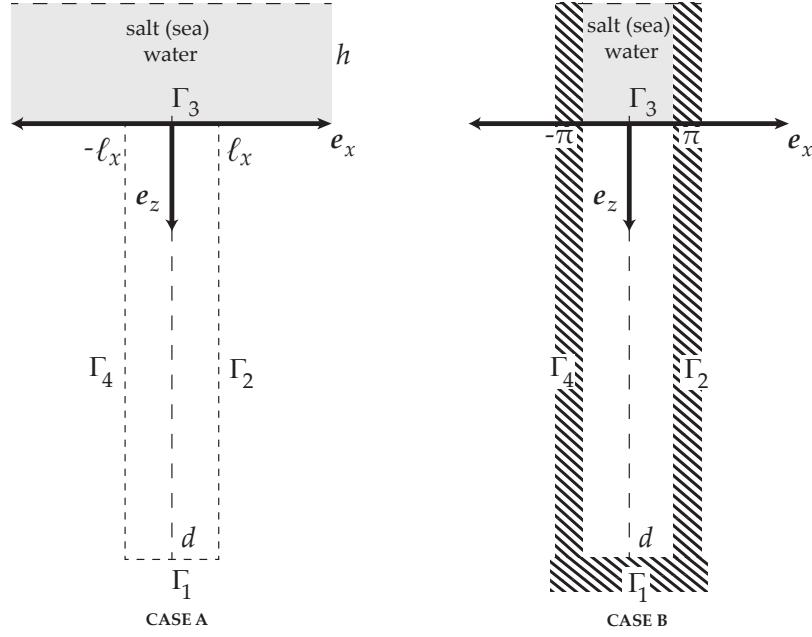


Fig. 2: Sketch of a two-dimensional cross-section of the flow domain in the x, z -plane. In the laterally unbounded **CASE A**: Periodic boundary conditions for the concentration c , the discharge \mathbf{u} , and pressure p apply at vertical boundary $\Gamma_v = \Gamma_2 \cup \Gamma_4$. In the laterally bounded **CASE B**: No-flow and no-transport conditions apply at the vertical boundary Γ_v . For numerical simulations, we consider a truncated column of depth d .

5.1. Linearized stability

Assuming the perturbations to be small, one disregards the nonlinear term $\mathbf{u} \cdot \nabla c$ in (25a). The result is the perturbation system

$$\frac{\partial c}{\partial t} + R w \frac{\partial C_0}{\partial z} - \Delta c = 0, \quad (39a)$$

$$\nabla \cdot \left(\frac{1}{\kappa(z)} \nabla w \right) = \Delta_{\perp} c, \quad (39b)$$

$$\frac{\partial w}{\partial z} = \kappa(z) \Delta_{\perp} p, \quad (39c)$$

in Ω_A or Ω_B , with

$$c = 0, \quad \frac{\partial w}{\partial z} = 0, \quad p = 0 \quad \text{at } z = 0, \text{ for all } t > 0, \quad (40a)$$

$$c \rightarrow 0, \quad w \rightarrow 0, \quad \frac{\partial p}{\partial z} \rightarrow 0 \quad \text{as } z \rightarrow \infty, \text{ for all } t > 0, \quad (40b)$$

and with

$$\text{(CASE A)} \quad \text{Periodic conditions at } \Gamma_v \quad (40c)$$

$$\text{(CASE B)} \quad \text{No-flow and no-transport conditions at } \Gamma_v \quad (40d)$$

Note that (39) is a linear system in terms of c , w (z -component of Darcy velocity), and p . Because of the linearity and the nature of the conditions (40a) and (40b), we may assume that solutions are of the form

$$\{c, w, p\}(x, y, z, t) = \{c, w, p\}(z, t) \cos a_x x \cos a_y y, \quad (41)$$

Then (40c), (40d) are satisfied provided the wavenumbers satisfy

$$\text{(CASE A)} \quad a_x = \frac{n\pi}{\ell_x}, \quad a_y = \frac{m\pi}{\ell_y} \quad n, m = 1, 2, \dots \quad (42a)$$

$$\text{(CASE B)} \quad a_x = n, \quad a_y = \frac{m}{\ell} \quad n, m = 1, 2, \dots \quad (42b)$$

Substituting (41) in (39) gives for the amplitudes $c(z, t)$, $w(z, t)$, and $p(z, t)$ the equations

$$\frac{\partial c}{\partial t} + R w \frac{\partial C_0}{\partial z} = \frac{\partial^2 c}{\partial z^2} - a^2 c, \quad (43a)$$

$$-\frac{\partial}{\partial z} \left(\frac{1}{\kappa(z)} \frac{\partial w}{\partial z} \right) + \frac{a^2}{\kappa(z)} w = a^2 c, \quad (43b)$$

$$\frac{\partial w}{\partial z} = -a^2 \kappa(z) p, \quad (43c)$$

for $z > 0$ and $t > 0$, subject to conditions (40a), (40b). Here $a^2 = a_x^2 + a_y^2$.

We will study the linearized problem (43) for all $a > 0$ and $R > 0$. However, when interpreting the results in terms of stability we consider

$$\text{(CASE A)} \quad a \text{ unspecified and } R \equiv 1$$

$$\text{(CASE B)} \quad a = a_{n,m} = \sqrt{n^2 + (m/\ell)^2} \text{ and } R = R_S,$$

where R_S denotes the system Rayleigh number as specified in Table 1.

It suffices to consider equations (43a) and (43b) for c and w only, since p follows directly from equation (43c).

In the following sections 5.1.1 and 5.1.2 we use two distinct methods to obtain linear stability results.

5.1.1 Direct L^2 -estimate

To obtain an estimate for the L^2 -norm of the concentration perturbation, we multiply equation (43a) by c and integrate the result in z . This gives the expression

$$\begin{aligned} \frac{1}{2} \frac{d}{dt} \int_0^\infty c^2(z, t) dz + \int_0^\infty \left(\frac{\partial c}{\partial z}(z, t) \right)^2 dz \\ = -a^2 \int_0^\infty c^2(z, t) dz - R \int_0^\infty c(z, t) w(z, t) \frac{\partial C_0}{\partial z}(z, t) dz \end{aligned} \quad (44)$$

for each $t > 0$.

Since the ground state saturation C_0 satisfies

$$0 < -\frac{\partial C_0}{\partial z}(z, t) \leq -\frac{\partial C_0}{\partial z}(0, t) = \frac{1}{\sqrt{\pi t}} \quad \text{for } z > 0 \text{ and } t > 0, \quad (45)$$

we can estimate (44) by, disregarding the second term in the left side as well,

$$\frac{1}{2} \frac{d}{dt} \int_0^\infty c^2(z, t) dz \leq -a^2 \int_0^\infty c^2(z, t) dz + \frac{R}{\sqrt{\pi t}} \int_0^\infty |c(z, t) w(z, t)| dz$$

for $t > 0$.

Introducing the L^2 -norm (i.e. $\|c\|_2^2 = \int_0^\infty c^2 dz$) and dropping t from the notation, this inequality is written in the compact form

$$\frac{1}{2} \frac{d}{dt} \|c\|_2^2 \leq -a^2 \|c\|_2^2 + \frac{R}{\sqrt{\pi t}} \int_0^\infty |c w| dz. \quad (46)$$

At this point we use the Cauchy–Schwarz inequality

$$\int_0^\infty |cw| \, dz \leq \|c\|_2 \|w\|_2 ,$$

to estimate (46) further as

$$\frac{1}{2} \frac{d}{dt} \|c\|_2^2 \leq -a^2 \|c\|_2^2 + \frac{R}{\sqrt{\pi t}} \|c\|_2 \|w\|_2 . \quad (47)$$

Next we multiply (43b) by w and integrate the result. Using the Neumann condition (28) this yields

$$\int_0^\infty \frac{1}{\kappa(z)} \left(\frac{\partial w}{\partial z} \right)^2 dz + a^2 \int_0^\infty \frac{1}{\kappa(z)} w^2 dz = a^2 \int_0^\infty wc \, dz . \quad (48)$$

Now let $\bar{\kappa} = \max_{z \geq 0} \kappa(z)$. Then we deduce from (48)

$$\begin{aligned} \frac{1}{\bar{\kappa}} \|w\|_2^2 &\leq \int_0^\infty |w| |c| \, dz \\ &\leq \|w\|_2 \|c\|_2 , \end{aligned}$$

where we used again Cauchy–Schwarz. Thus

$$\|w(t)\|_2 \leq \bar{\kappa} \|c(t)\|_2 \quad \text{for } t > 0 . \quad (49)$$

Applying this inequality in (47) gives

$$\frac{d}{dt} \|c\|_2^2 \leq \left(-2a^2 + \frac{2R\bar{\kappa}}{\sqrt{\pi t}} \right) \|c\|_2^2 . \quad (50)$$

Integration this expression in t yields the estimate

$$\|c(t)\|_2 \leq \|c(0)\|_2 e^{-a^2 t + \frac{2R\bar{\kappa}}{\sqrt{\pi}} \sqrt{t}} . \quad (51)$$

By (49), the same holds for $\|w(t)\|_2$. Next we apply (49) in (48) to find

$$\int_0^\infty \frac{1}{\kappa(z)} \left(\frac{\partial w}{\partial z} \right)^2 dz \leq a^2 \int_0^\infty |wc| \, dz \leq a^2 \|w\|_2 \|c\|_2 \leq a^2 \bar{\kappa} \|c\|_2^2 .$$

Using this inequality and (43c) we obtain

$$\int_0^\infty \kappa(z) p^2 \, dz = \frac{1}{a^4} \int_0^\infty \frac{1}{\kappa(z)} \left(\frac{\partial w}{\partial z} \right)^2 dz \leq \frac{\bar{\kappa}}{a^2} \|c\|_2^2 .$$

Hence

$$\|p(t)\|_2 \leq \frac{1}{a} \sqrt{\frac{\bar{\kappa}}{\underline{\kappa}}} \|c(t)\|_2 \quad \text{for } t > 0 , \quad (52)$$

where $\underline{\kappa} = \min_{z \geq 0} \kappa(z)$.

Thus both the pressure p and the vertical flow w decay like c . Estimate (51) says that x, y -periodic perturbations, characterized by the wave number a , eventually decay in time if the initial

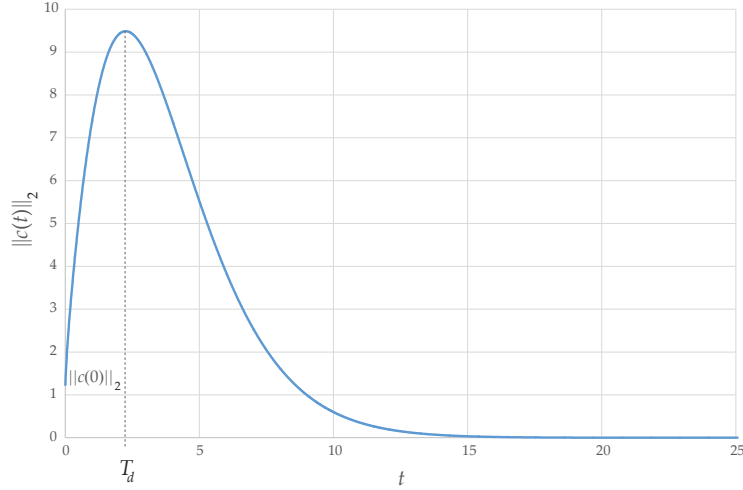


Fig. 3: Stability upperbound by direct L^2 -estimate: right-hand side in (51).

amplitude $c(0) \in L^2([0, \infty))$ (i.e. $\int_0^\infty c(0)^2 dz < \infty$). This is due to the term $-a^2 t$ in the exponent. In essence, this is a diffusive term. This can be seen when going back to the dimensional form

$$a^2 t \longrightarrow \begin{cases} \left(a \frac{D}{Q_c}\right)^2 \frac{Q_c^2 t}{\phi D}, & \text{(CASE A)} \\ (a L_c)^2 \frac{D t}{\phi L_c^2}, & \text{(CASE B)} \end{cases}$$

where now a and t are dimensional. A sketch of the upperbound in (51) is shown in Figure 3. Note that the maximum is reached at

$$t = T_d = \frac{R^2 \bar{\kappa}^2}{a^4 \pi}.$$

The stability result expressed by (51) was derived using the rather crude estimate (45). In the next section we do not use (45). Instead we consider the growth or decay of small perturbations of $C_0(z, t)$ for each $t > 0$.

For future reference we introduce

$$R_M = R_M(a, t) := \frac{a^2 \sqrt{\pi t}}{\bar{\kappa}}. \quad (53)$$

Using this in (50) gives the form

$$\frac{d}{dt} \|c\|_2^2 \leq -2a^2 \left\{ 1 - \frac{R}{R_M} \right\} \|c\|_2^2, \quad (54)$$

implying that if $R < R_M(a, t)$, then L^2 -stability of the linearized system is guaranteed.

5.1.2 Eigenvalue formulation

Again we start from the linear system (43). In case of a steady, i.e. time independent, ground state, as in Wooding et al. (1997a,b), one looks for amplitudes c and w that have an exponential growth rate in time, i.e. $\{c, w\}(z, t) = \{c, w\}(z) e^{\sigma t}$. If $\sigma < 0$, then all small perturbations decay in time and we call the system linearly asymptotic stable. If, on the other hand, $\sigma > 0$ then small perturbations grow exponentially in time and we call the system unstable.

In this paper the ground state depends on time as well. However, since $\frac{\partial C_0}{\partial z}(z, t + \tau) \approx \frac{\partial C_0}{\partial z}(z, t)$ for $|\tau|$ being sufficiently small, we proceed similarly as in the steady case. This is called the ‘frozen-profile’ approach (van Duijn et al., 2002).

Hence, for given $t > 0$, we consider instead of (43a) the approximate equation

$$\frac{\partial c}{\partial \tau} + R w \frac{\partial C_0}{\partial z}(z, t) = \frac{\partial^2 c}{\partial z^2} - a^2 c, \quad (55)$$

for $\tau > 0$, but sufficiently small. Now time t appears as a parameter in the equation. We then assume that the perturbations have an exponential growth rate denoted by σ , i.e.

$$\{c, w\}(z, \tau) = \{c, w\}(z) \exp(\sigma \tau). \quad (56)$$

Substituting (56) into (55) and (43b) yields the following eigenvalue problem: For given $a > 0$, $t > 0$ and σ , find the smallest $R > 0$ (denoted by $R = R_1(a, t, \sigma)$) so that the problem

$$\sigma c = \frac{d^2 c}{dz^2} - a^2 c - R w \frac{\partial C_0}{\partial z}(z, t), \quad (57a)$$

$$-\frac{d}{dz} \left(\frac{1}{\kappa(z)} \frac{dw}{dz} \right) + \frac{a^2}{\kappa(z)} w = a^2 c, \quad (57b)$$

for $z > 0$, with

$$c = 0, \quad \frac{dw}{dz} = 0 \quad \text{at } z = 0, \quad (57c)$$

$$c \rightarrow 0, \quad w \rightarrow 0 \quad \text{as } z \rightarrow \infty \quad (57d)$$

has a non-trivial solution.

The eigenvalues now depend on a , t , and on the growth rate σ . In the salt lake problem (Pieters and van Duijn, 2006) it was shown that the smallest positive eigenvalue $R_1(a, t, \sigma)$ satisfies

$$R_1(a, t, \sigma) \leq R_1(a, t, 0) \quad \text{if and only if } \sigma \leq 0. \quad (58)$$

This expresses exchange of stability. We assume here that (58) also holds for the eigenvalue problem (57). Under this assumption, which is supported by numerical evidence in Appendix C, it suffices to analyse (57) for the case of neutral stability, i.e. $\sigma \equiv 0$.

For completeness we write the eigenvalue problem resulting from linearized stability.

$$(P_L) \left\{ \begin{array}{l} \text{Find smallest } R > 0 \text{ so that the system} \\ \left. \begin{array}{l} \frac{d^2 c}{dz^2} - a^2 c - R w \frac{\partial C_0}{\partial z}(z, t) = 0 \\ -\frac{d}{dz} \left(\frac{1}{\kappa(z)} \frac{dw}{dz} \right) + \frac{a^2}{\kappa(z)} w = a^2 c \end{array} \right\} \quad \text{for } z > 0, \\ \text{with } c \text{ and } w \text{ satisfying} \\ \left. \begin{array}{l} c = 0, \quad \frac{dw}{dz} = 0 \\ c \rightarrow 0, \quad w \rightarrow 0 \end{array} \right\} \quad \begin{array}{l} \text{at } z = 0, \\ \text{as } z \rightarrow \infty, \end{array} \\ \text{has a nontrivial solution.} \end{array} \right. \quad (59)$$

Note that here $a > 0$ and $t > 0$ act as parameters. The smallest positive eigenvalue is denoted by $R_L = R_L(a, t)$. We return to this problem in section 6.1.1.

The behaviour of large perturbations is studied in the next section.

5.2. The variational energy method

The method of linearised stability holds for infinitesimal small perturbations. What can be said about the system behavior for larger perturbations. To address this point, we apply the variational energy approach, see van Duijn et al. (2002) for the specific application to the salt lake problem and Straughan (1992, 2008) for a general overview.

Remark 3 *The energy method can be applied for both domain Ω_A (CASE A) and Ω_B (CASE B). In the derivation that follows, we use Ω_i to denote the flow domain. The index i can refer to A or B. Both geometrical cases would give the same equations, although the interpretation of the stability results is different.*

In the energy method one estimates the time derivative of the L^2 -norm of the concentration perturbation c . In particular, the aim is to find the largest R -interval for which

$$\frac{d}{dt} \int_{\Omega_i} c^2 < 0 \quad \text{for all } t > 0. \quad (60)$$

Here and in the integrals below we disregard the infinitesimal volume elements in the notation. The related maximum R -value clearly will depend on the wavenumber a and, because $C_0 = C_0(z, t)$, on time t . Starting point here is the nonlinear system (25).

To investigate (60), we multiply (25a) by c and integrate over Ω_i . Using (25c) and the boundary conditions along Γ_v , we find the identity:

$$\frac{1}{2} \frac{d}{dt} \int_{\Omega_i} c^2 = - \int_{\Omega_i} |\nabla c|^2 - R \int_{\Omega_i} cw \frac{\partial C_0}{\partial z}. \quad (61)$$

Thus, if R is chosen such that the right-hand side of (61) is negative for all perturbations satisfying a given constraint which is closest to the solution of the problem, then stability is guaranteed.

Roughly, there exist two approaches to incorporate the constraint on the admissible perturbations. One is based on (25c) (see Homsy and Sherwood (1976)), and the integrated Darcy equation (25b):

$$\nabla \cdot \mathbf{u} = 0, \quad \int_{\Omega_i} \frac{1}{\kappa(z)} |\mathbf{u}|^2 - \int_{\Omega_i} cw = 0. \quad (62)$$

The other approach, and the one we focus on in this paper, uses the pointwise differential expression (31) as constraint:

$$\nabla \cdot \left(\frac{1}{\kappa(z)} \nabla w \right) = \Delta_{\perp} c \quad \text{in } \Omega_i.$$

Applied to the salt lake problem, van Duijn et al. (2002) showed that this gives sharper results than constraint (62). It leads to the following class of admissible perturbations:

$$\mathcal{H} = \left\{ (c, w) \neq (0, 0) : c, w \text{ (or } \mathbf{u}) \text{ are } x, y\text{-periodic with respect to } \Omega_i, \right. \\ \left. c = \frac{\partial w}{\partial z} = 0 \text{ at } z=0, c=w=0 \text{ as } z \rightarrow \infty, \text{ and } \nabla \cdot \left(\frac{1}{\kappa(z)} \nabla w \right) = \Delta_{\perp} c \text{ in } \Omega_i \right\}.$$

In the energy method the goal is to determine the largest R -interval for which

$$\int_{\Omega_i} |\nabla c|^2 + R \int_{\Omega_i} cw \frac{\partial C_0}{\partial z} > 0 \quad \text{for all } (c, w) \in \mathcal{H}.$$

To this end we consider the maximum problem

$$\frac{1}{R_E} = \sup_{(c,w) \in \mathcal{H}} \frac{- \int_{\Omega_i} cw \frac{\partial C_0}{\partial z}}{\int_{\Omega_i} |\nabla c|^2} . \quad (63)$$

Suppose this maximum problem has a solution (i.e. there exists such a $R_E \in (0, \infty)$). Then any solution of (25), (26), and (28) satisfies

$$\frac{1}{R_E} \geq \frac{- \int_{\Omega_i} cw \frac{\partial C_0}{\partial z}}{\int_{\Omega_i} |\nabla c|^2} ,$$

and thus for any $R < R_E$ one obtains from (61)

$$\frac{1}{2} \frac{d}{dt} \int_{\Omega_i} c^2 \leq - \left(1 - \frac{R}{R_E}\right) \int_{\Omega_i} |\nabla c|^2 < 0 .$$

Hence for $R < R_E$ the system is unconditionally stable.

Maximum problem (63) is equivalent to the following minimum problem: Find the smallest $R_E > 0$ so that

$$\inf_{(c,w) \in \mathcal{H}} \left\{ \int_{\Omega_i} |\nabla c|^2 + R_E \int_{\Omega_i} cw \frac{\partial C_0}{\partial z} \right\} = 0 . \quad (64)$$

To solve (64), we introduce the functional

$$J(c, w, \pi; R) = \int_{\Omega_i} |\nabla c|^2 + R \int_{\Omega_i} cw \frac{\partial C_0}{\partial z} + \int_{\Omega_i} \pi \left[\nabla \cdot \left(\frac{1}{\kappa(z)} \nabla w \right) - \Delta_{\perp} c \right] , \quad (65)$$

where π acts as a Lagrange multiplier. We need to find points (c, w, π) , with $(c, w) \in \mathcal{H}$ and π is Ω_i -periodic, $\lim_{z \rightarrow \infty} \pi(z) = 0$, and values of R where the first variation of (65) vanishes. This leads to the Euler–Lagrange equations

$$-2\Delta c + R \frac{\partial C_0}{\partial z} w = \Delta_{\perp} \pi , \quad (66a)$$

$$\nabla \cdot \left(\frac{1}{\kappa(z)} \nabla \pi \right) = -R \frac{\partial C_0}{\partial z} c , \quad (66b)$$

$$\nabla \cdot \left(\frac{1}{\kappa(z)} \nabla w \right) = \Delta_{\perp} c , \quad (66c)$$

in Ω_i . Moreover, $\frac{\partial \pi}{\partial z}(0) = 0$ appears as a natural boundary condition. Details of the derivation are given in Appendix D.

Thus far Ω_i has been left general. Explicitly considering periodicity of the form (41) for c, π ,

and w , which is allowed since the equations in (66) are linear, yields the eigenvalue problem

$$\begin{aligned}
 (P_E) \left\{ \begin{array}{l} \text{Find the smallest } R > 0 \text{ so that the system} \\ \left. \begin{array}{l} \frac{d^2 c}{dz^2} - a^2 c = \frac{1}{2} R \frac{\partial C_0}{\partial z} w + \frac{1}{2} a^2 \pi \\ -\frac{d}{dz} \left(\frac{1}{\kappa(z)} \frac{d\pi}{dz} \right) + \frac{a^2}{\kappa(z)} \pi = R \frac{\partial C_0}{\partial z} c \\ -\frac{d}{dz} \left(\frac{1}{\kappa(z)} \frac{dw}{dz} \right) + \frac{a^2}{\kappa(z)} w = a^2 c \end{array} \right\} & \text{for } z > 0, \quad (67a) \\ \\ \text{with } c, w, \text{ and } \pi \text{ satisfying} \\ c = \frac{dw}{dz} = \frac{d\pi}{dz} = 0 & \text{at } z = 0, \quad (67b) \\ c \rightarrow 0, \quad w \rightarrow 0, \quad \pi \rightarrow 0 & \text{as } z \rightarrow \infty, \quad (67c) \\ \text{has a nontrivial solution.} \end{array} \right.
 \end{aligned}$$

Here the smallest positive eigenvalue is denoted by $R_E = R_E(a, t)$. We return to this problem in Section 6.

6. STABILITY RESULTS

6.1. Homogeneous soils

When considering homogeneous soils, we set $\kappa(z) = 1$ in problems (P_L) and (P_E) . We show below that this makes it possible to eliminate the time t from the eigenvalue problems. As a consequence we obtain a clear interpretation of the role of a , R_S and t in the stability behaviour.

6.1.1 Eigenvalue approach

Since

$$\frac{\partial C_0(z, t)}{\partial z} = -\frac{1}{\sqrt{\pi t}} e^{-\frac{z^2}{4t}},$$

we can remove t from (P_L) and (P_E) by introducing the variables

$$\eta := \frac{z}{\sqrt{t}}, \quad \{c, w, \pi\}(\eta) := \{c, w, \pi\}(z, t), \quad b := a\sqrt{t}, \quad \text{and } R^* := \sqrt{\frac{t}{\pi}} R. \quad (68)$$

This yields the transformed problems

$$(P_L^*) \left\{ \begin{array}{l} \text{Find smallest } R > 0 \text{ so that the system} \\ \left. \begin{array}{l} -\frac{d^2 c}{d\eta^2} + b^2 c = R^* w e^{-\eta^2/4} \\ -\frac{d^2 w}{d\eta^2} + b^2 w = b^2 c \end{array} \right\} \quad \text{for } \eta > 0, \\ \text{with } c \text{ and } w \text{ satisfying} \\ c = 0, \quad \frac{dw}{d\eta} = 0 \quad \text{at } \eta = 0, \\ c \rightarrow 0, \quad w \rightarrow 0 \quad \text{as } \eta \rightarrow \infty, \\ \text{has a nontrivial solution,} \end{array} \right. \quad (69)$$

and

$$(P_E^*) \left\{ \begin{array}{l} \text{Find smallest } R > 0 \text{ so that the system} \\ \left. \begin{array}{l} -\frac{d^2 c}{d\eta^2} + b^2 c = \frac{R^*}{2} w e^{-\eta^2/4} - \frac{b^2}{2} \pi \\ -\frac{d^2 \pi}{d\eta^2} + b^2 \pi = -R^* c e^{-\eta^2/4} \\ -\frac{d^2 w}{d\eta^2} + b^2 w = b^2 c \end{array} \right\} \quad \text{for } \eta > 0, \\ \text{with } c, w, \text{ and } \pi \text{ satisfying} \\ c = \frac{dw}{d\eta} = \frac{d\pi}{d\eta} = 0 \quad \text{at } \eta = 0, \\ c \rightarrow 0, \quad w \rightarrow 0, \quad \pi \rightarrow 0 \quad \text{as } \eta \rightarrow \infty. \\ \text{has a nontrivial solution.} \end{array} \right. \quad (70)$$

Hence, the eigenvalue problems are written in terms of the same similarity variable as the ground state $C_0(z, t)$.

To find the smallest possible eigenvalues for $b > 0$, we solve (P_L^*) and (P_E^*) by means of a Chebyshev–Petrov–Galerkin spectral method. Details and references are given in Appendix H. The result is shown in Figure 4. Note that $R_L^*(b) > R_E^*(b)$ for all $b > 0$, but the difference is small for $b > 0.5$. This is different from the salt lake problem (van Duijn et al. (2002)), where there is a significant gap between R_L and R_E . In this gap small perturbations vanish, but large perturbations may grow in time. Since for all practical purposes, see Figure 4,

$$R_L^*(b) \approx R_E^*(b) \quad \text{for } b > 0, \text{ say,} \quad (71)$$

the ponded problem considered in this paper has a sharp stability bound: below R_E^* there is unconditional stability and above R_L^* instability.

Remark 4 For **CASE A** there is no defined system Rayleigh number R_S . However, one still has the Rayleigh number R^* showing up in the eigenvalue problems due to the similarity transformation. This system Rayleigh number then refers to time only, as $R^* = \sqrt{\frac{t}{\pi}}$.

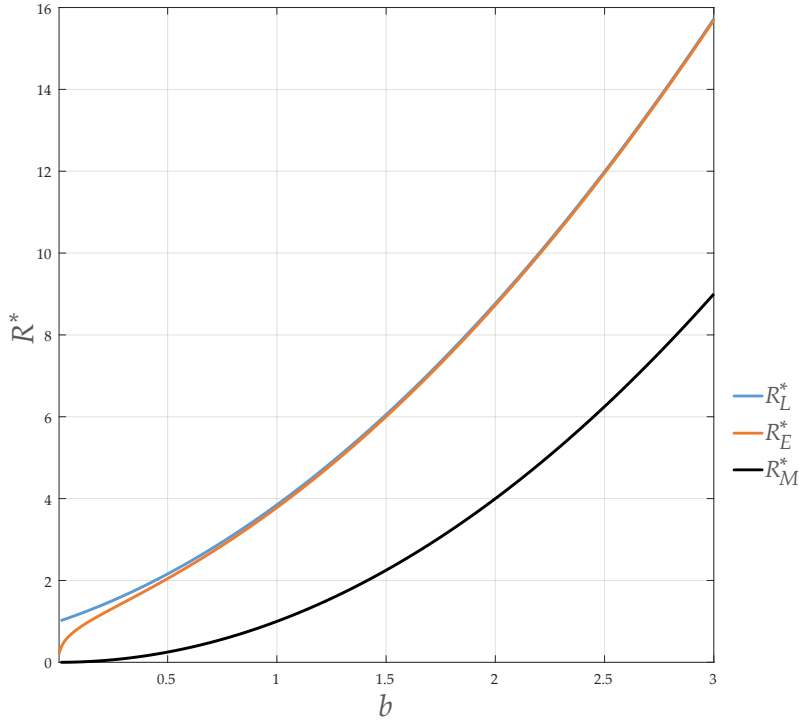


Fig. 4: Stability curves for the ponded case, according to the direct approach R_M^* , method of linearised stability R_L^* and energy method with differential constraint R_E^* .

Remark 5 Using Green's functions, one can write w and π in (P_L^*) and (P_E^*) in terms of c . Substituting this into (69) and (70), and applying again the Green's functions, yields eigenvalue problems in terms of integral equations for c . These transformations and some mathematical properties of the integral forms are detailed in Appendix F. In particular, since $c(0) = \frac{dw}{d\eta}(0) = \frac{d\pi}{d\eta}(0) = 0$, the corresponding integral equations are non-selfadjoint.

We show in Appendix E that both $R_L^*(b)$ and $R_E^*(b)$ have the lower bound:

$$R_L^*(b), R_E^*(b) > b^2 \quad \text{for all } b > 0.$$

This estimate is obtained by manipulating the equations in (P_L^*) and (P_E^*) . Recalling that, see (53) with $\bar{\kappa} = 1$, $R_M = R_M(a, t) = a^2 \sqrt{\pi t}$, we find after applying (68),

$$R_M^*(b) = \sqrt{\frac{t}{\pi}} R_M = b^2.$$

Hence the stability bound obtained by the direct L^2 -estimate of the linearized equations is a lower bound for both $R_L^*(b)$ and $R_E^*(b)$.

6.1.2 Stability interpretation

How to interpret and estimate the (in)stability from Figure 4? For simplicity we assume that (71) holds. The fact that $R_L^*(b) > R_E^*(b)$ close to $b = 0$ requires only a minor modification of the argument. Let $a > 0$ be a given wavenumber and let R_S be the system Rayleigh number. In view

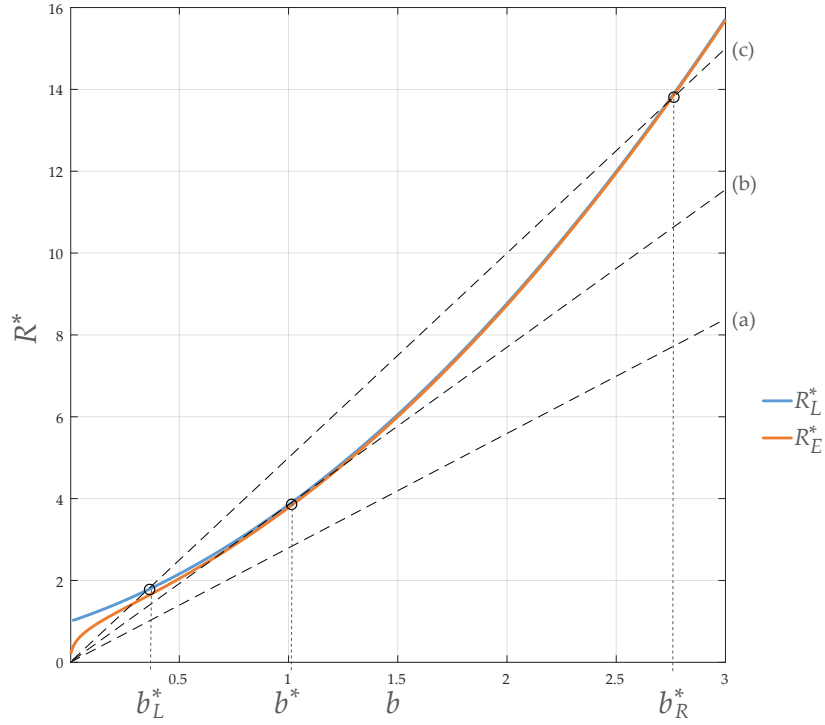


Fig. 5: Repetition of Figure 4, but now it includes dashed lines $R_S^*(b) = \frac{R_S}{a\sqrt{\pi}}b$ for three different pairs (a, R_S) : (a) case where $R^*(b)$ lies below $R_E^*(b)$ - system unconditional stable for chosen (a, R_S) value; (b) border case where $R^*(b)$ touches $R_E^*(b)$ at $b=b^*$; (c) case where $R^*(b)$ intersects $R_L^*(b)$ in b_L^* and b_R^* .

of (68) we have

$$\begin{aligned}
 \text{CASE A} \quad R_S^* &= \sqrt{\frac{t}{\pi}} = \frac{b}{a\sqrt{\pi}} =: R_S^*(b) . \\
 \text{CASE B} \quad R_S^* &= \sqrt{\frac{t}{\pi}} R_S = \frac{R_S}{a\sqrt{\pi}} b =: R_S^*(b) .
 \end{aligned} \tag{72}$$

Hence $R_S^*(b)$ is a straight line passing through the origin in the (b, R^*) -plane. With reference to Figure 5, three cases can be distinguished, depending on whether the straight line lies (a) entirely below R_E^* , or (b) touches R_E^* at b^* , or (c) intersects R_E^* at b_L^* and b_R^* . We successively consider cases (b), (a), and (c).

Let $\alpha > 0$ be such that the line ab is below $R_E^*(b)$, except at $b=b^*$ where it touches R_E^* :

$$ab^* = R_E^*(b^*) \quad \text{and} \quad ab < R_E^*(b) \quad \text{for } b \neq b^* . \tag{73}$$

Numerically we found

$$\alpha = \alpha_{\text{crit}} \approx \frac{6.82}{\sqrt{\pi}} \approx 3.85 \quad \text{and} \quad b = b^* \approx 1.01 .$$

Writing $R^*(b) = ab$, we have for $\alpha < \alpha_{\text{crit}}$

$$R^*(b) < R_E^*(b) \quad \text{for all } b > 0 .$$

Thus if

$$\begin{aligned} \text{CASE A} \quad & \frac{1}{a\sqrt{\pi}} < \alpha_{\text{crit}} \quad \text{or} \quad a > \frac{1}{6.82}, \\ \text{CASE B} \quad & \frac{R_S}{a\sqrt{\pi}} < \alpha_{\text{crit}} \quad \text{or} \quad a > \frac{R_S}{6.82}, \end{aligned}$$

then the ground state is unconditional stable.

How to interpret the result?

CASE A: When considering a specific horizontal structure, for certain ℓ_x and ℓ_y , the wavenumber a has the discrete values

$$a_{n,m} = \sqrt{\left(\frac{n\pi}{\ell_x}\right)^2 + \left(\frac{m\pi}{\ell_y}\right)^2} \quad n, m = 1, 2, \dots$$

Thus if the smallest value $a_{1,1}$ satisfies

$$a_{1,1} = \sqrt{\left(\frac{\pi}{\ell_x}\right)^2 + \left(\frac{\pi}{\ell_y}\right)^2} \geq \frac{1}{6.82}, \quad (74)$$

then the ground state is L^2 -stable for perturbations that are x, y -periodic with respect to a rectangular structure with ℓ_x and ℓ_y satisfying (74). But perturbations with larger wavelengths (or smaller wavenumbers) may grow in time to form finger-like patterns.

CASE B: Then a takes the discrete values

$$a_{n,m} = \sqrt{n^2 + \left(\frac{m}{\ell}\right)^2} \quad n, m = 1, 2, \dots$$

If the horizontal dimensions of the vertical column Ω_B satisfy

$$a_{1,1} = \sqrt{1 + \left(\frac{1}{\ell}\right)^2} \geq \frac{R_S}{6.82}, \quad (75)$$

then the ground state is L^2 -stable in Ω_B implying that no fingers can occur as a result of initial perturbations.

If $\alpha > \alpha_{\text{crit}}$, then $R_S^*(b)$ intersects $R_L^*(b) \approx R_E^*(b)$ at the points b_L^* and b_R^* as indicated in Figure 5:

$$R_S^*(b) > R_E^*(b) \quad \text{for } b_L^* < b < b_R^* \quad \text{and} \quad R_S^*(b) < R_E^*(b) \quad \text{elsewhere}.$$

Clearly the intersection points b_L^* and b_R^* depend on a and R_S : $b_i^* = b_i^*(a, R_S)$ for $i = L, R$. For a given system Rayleigh number R_S , the boundary curves in the (a, t) -plane corresponding to $a\sqrt{t} = b_i^*(a, R_S)$ ($i = L, R$) are shown in Figure 6 (Left: **CASE A**, Right: **CASE B**). These curves meet when $\frac{R_S}{a} = 6.82$, since then $b_L^* = b_R^* = b^*$.

Above the composite curve (red and blue curves in Figure 5) one has L^2 -stability and below the curve small perturbations grow in time. Let $a = a_{1,1}$, see (74) or (75), denote again the smallest wavenumber for a given Ω_A or Ω_B : i.e. the most unstable mode for a given geometry. Then, with reference to Figure 6, perturbations decay in time when $0 < t < t_L(a)$, they grow when $t_L(a) < t < t_R(a)$, and finally again decay when $t > t_R(a)$. Here $t_i(a) = (b_i^*/a)^2$ with $i = L, R$.

In Section 7 we return to Figures 5 and 6 for the interpretation in terms of the physical parameters.

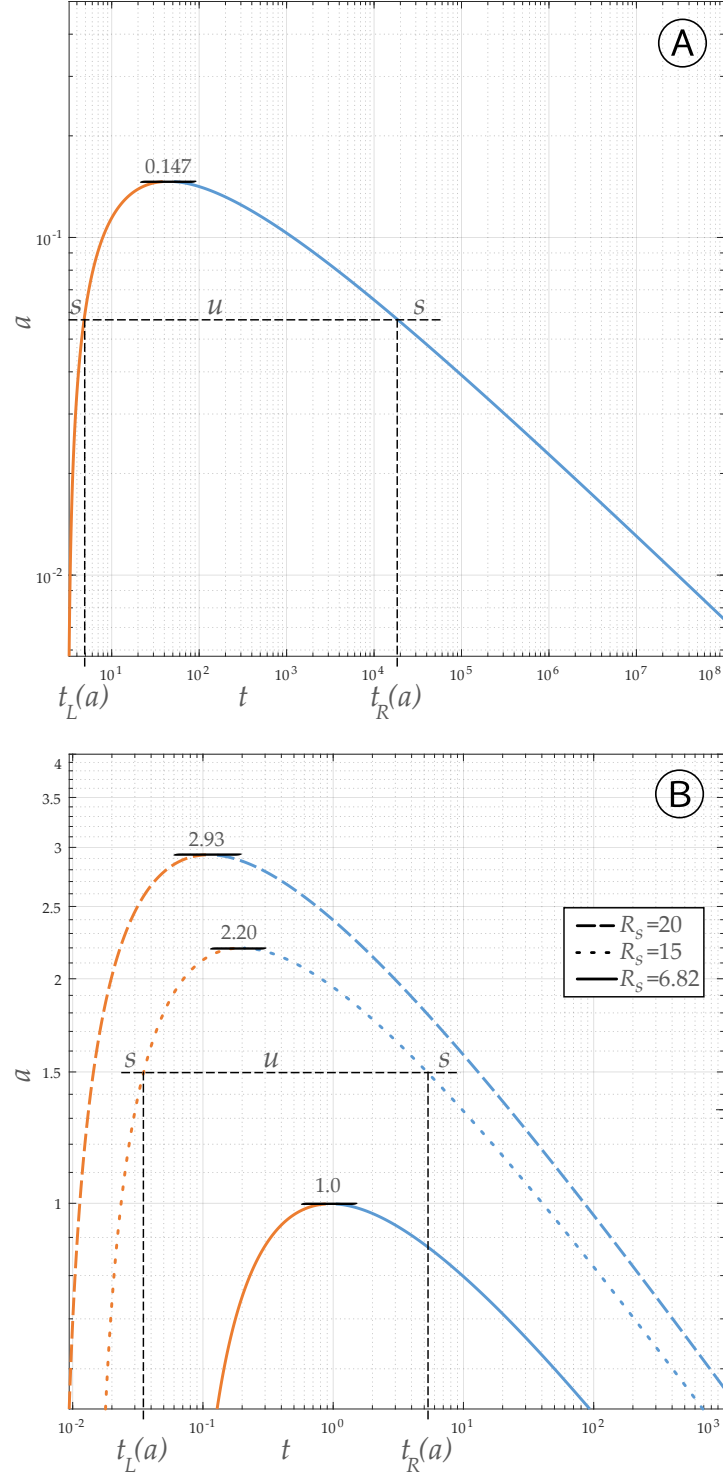


Fig. 6: Regions of (in)stability in (t, a) -plane. The blue curves correspond to $t = (b_R^*/a)^2$, and the red curves to $t = (b_L^*/a)^2$. (A): **CASE A**. (B): **CASE B** for $R_S = 20$, $R_S = 15$, and $R_S = 6.82$. For the maxima it holds that $a = b^*/\sqrt{t}$. For $t < t_L(a)$ and $t > t_R(a)$ the system is stable (indicated by s), while for $t_L(a) < t < t_R(a)$ the system has an episode of transient instability (indicated by u): perturbations having wave number a will first decay, then grow and eventually decay again in time.

6.1.3 Direct numerical simulations

To verify and interpret our theoretical findings we make a comparison with the nonlinear dynamics described by the full set of equations (14). To this end we perform direct numerical simulations of (14) in two-dimensional cells:

$$\begin{aligned} \text{CASE A} \quad \widehat{\Omega}_A^d &:= \Omega_A^d \cap \{y=0\} ; \\ \text{CASE B} \quad \widehat{\Omega}_B^d &:= \Omega_B^d \cap \{y=0\} , \end{aligned}$$

see also Figure 2.

Having a two-dimensional flow domain in which $\nabla \cdot \mathbf{U} = 0$, allows us to formulate the convective flow part in (14) in terms of the stream function Ψ ; i.e.

$$\mathbf{U} = (q_x, q_z) \quad \text{where} \quad q_x = -\frac{\partial \Psi}{\partial z}, \quad q_z = \frac{\partial \Psi}{\partial x}. \quad (76)$$

Taking the two-dimensional curl of the Darcy equation (14b) and substituting (76) into the result gives for Ψ the equation (de Josselin de Jong, 1960; Bear, 1972)

$$\nabla \cdot \left(\frac{1}{\kappa(z)} \nabla \Psi \right) = \frac{\partial C}{\partial x}. \quad (77)$$

In the computations presented in this Section we use $\kappa(z) = 1$, since the soil is considered homogeneous. Later, in Section 6.2, we allow $\kappa = \kappa(z)$.

The salt transport equation in terms of Ψ becomes

$$\frac{\partial C}{\partial t} + R \left(\frac{\partial \Psi}{\partial x} \frac{\partial C}{\partial z} - \frac{\partial \Psi}{\partial z} \frac{\partial C}{\partial x} \right) = \Delta C. \quad (78)$$

At the top Γ_3 we have $P = 1$. This means that $\frac{1}{\kappa(0)} q_x = -\frac{\partial P}{\partial x} = 0$ at Γ_3 . Hence from (76)

$$C = 1 \quad \text{and} \quad \frac{\partial \Psi}{\partial z} = 0 \quad \text{at } \Gamma_3. \quad (79)$$

At the bottom Γ_1 we prescribe $C = 0$ and $\frac{\partial P}{\partial z} = 0$, see (14c). This means that $\frac{1}{\kappa(d)} q_z = 0$ at Γ_1 . Hence, again from (76),

$$C = 0 \quad \text{and} \quad \Psi = \text{constant} = 0 \quad \text{at } \Gamma_1. \quad (80)$$

For **CASE A**, we impose periodic boundary conditions on C and Ψ along the vertical boundary Γ_v . This means that for all $t > 0$

$$C|_{\Gamma_2} = C|_{\Gamma_4}, \quad \frac{\partial C}{\partial x}|_{\Gamma_2} = \frac{\partial C}{\partial x}|_{\Gamma_4}, \quad (81)$$

and

$$\Psi|_{\Gamma_2} = \Psi|_{\Gamma_4}, \quad \frac{\partial \Psi}{\partial x}|_{\Gamma_2} = \frac{\partial \Psi}{\partial x}|_{\Gamma_4}. \quad (82)$$

For **CASE B** we impose no-flow and no-transport conditions along Γ_v , i.e. for $t > 0$

$$\frac{\partial C}{\partial x}|_{\Gamma_2} = \frac{\partial C}{\partial x}|_{\Gamma_4} = 0, \quad (83)$$

and

$$\frac{\partial \Psi}{\partial z}|_{\Gamma_2} = \frac{\partial \Psi}{\partial z}|_{\Gamma_4} = 0. \quad (84)$$

We prescribe an initial concentration

$$C(x, z, 0) = C_i(x, z) \quad (85)$$

in $\widehat{\Omega}_A^d$ for **CASE A** or $\widehat{\Omega}_B^d$ for **CASE B**. Equations (77) and (78) and conditions (79)–(85) define the problem (NSP), the Numerical–Solution–Problem. This problem has a unique solution $\{C, \Psi\}$ for any reasonable initial concentration (85), e.g. see Clément et al. (1992).

Remark 6 Repeating $\widehat{\Omega}_A^d$ in horizontal direction yields a smooth $2\ell_x$ -periodic solution in the strip $\Omega^d := \{0 < z < d\}$. If the initial concentration C_i is **symmetric** with respect to $x=0$, then the same holds for the evolving concentration C and anti-symmetry holds for the corresponding stream function Ψ : i.e.

$$\left. \begin{aligned} C(x, z, t) &= C(-x, z, t) , \\ \Psi(x, z, t) &= -\Psi(-x, z, t) \end{aligned} \right\} \text{ for } (x, z) \in \widehat{\Omega}_A^d \text{ and } t > 0 .$$

In view of the periodic boundary conditions this implies

$$\left. \frac{\partial C}{\partial x} \right|_{\Gamma_2} = \left. \frac{\partial C}{\partial x} \right|_{\Gamma_4} = 0$$

and

$$\Psi|_{\Gamma_2} = \Psi|_{\Gamma_4} = 0$$

for all $t > 0$. Thus for symmetric initial data, there is no transport across the vertical boundaries. Clearly, this does not hold for arbitrary initial data.

Problem (NSP) is solved by means of a characteristic–Galerkin finite element method. In the method, the convective operator is split from the diffusive one and is written in the form of the material derivative. A detailed description of this method is given in Appendix I.

Let the flow domain $\widehat{\Omega}_i^d$, with $i = A, B$, be discretized by a set of elements, forming a mesh \mathcal{M} . The element size of the mesh \mathcal{M} will be denoted by $\Delta\mathcal{M}$. Further, time is discretized by the time-grid $t^n := n\Delta t$, $n = 1, 2, \dots, N$. We denote the numerical solution of problem (NSP) by $\{\widetilde{C}_m^n, \widetilde{\Psi}_m^n\}$, where m indexes over the elements in \mathcal{M} , and n indexes over a discrete time grid.

To quantify the temporal behavior of the full system subject to an initial perturbation, we consider as a stability measure the discrete L^2 -norm in Ω_i :

$$E(t^n) := \sqrt{\sum_{m \in \mathcal{M}} w_m [\widetilde{C}_m^n - \widetilde{C}_{0,m}^n]^2} .$$

In this definition, we do not use the analytical ground state C_0 evaluated per element and time-grid point, but rather its numerical approximation $\widetilde{C}_{0,m}^n$.

We discuss (direct) numerical simulations for both **CASE A** and **CASE B**. For **CASE A**, an overview of the stability results is shown in Table 2. There we present for different wavenumbers a stability overview based on the eigenvalue approach (Figure 6 (A)) and on direct numerical simulations using the quantity $E(t)$ (Figure 7 (A)). The stability behavior in Table 2 is derived from Figure 6 (A) and Figure 7 (A). Figure 8 shows direct numerical simulations of the salt concentration for the cases 6, 7, and 8. In those simulations we have used $\Delta\mathcal{M} = \pi/5 \approx 0.6$, $\Delta t = 0.05$, and $N = 50000$ which corresponds to a (dimensionless) time horizon of 2500. Further, we perturb the initial concentration by a small symmetric (with respect to $x=0$) perturbation of the form

$$C(x, z, 0) = \epsilon \cos a_{1,1} x , \quad (86)$$

Table 2: Overview of the selected geometries for **CASE A** and their stability (S = stable, U = unstable episode).

#	ℓ_x	d	$a_{1,1}$	Figure 6 (A)	Figure 7 (A)
1	5π	250	0.2	S	S
2	6π	250	0.167	S	S
3	7π	250	0.143	S	S
4	8π	250	0.125	U	U
5	9π	250	0.111	U	U
6	10π	250	0.1	U	U
7	11π	250	0.0909	U	U
8	12π	250	0.0833	U	U

Table 3: Overview of the selected system Rayleigh numbers for **CASE B** and their stability (S = stable, U = unstable episode).

#	R_S	d	$a_{1,1}$	Figure 6 (B)	Figure 7 (B)
1	1	50	1	S	S
2	6	50	1	S	S
3	8	50	1	U	U
4	9	50	1	U	U
5	10	50	1	U	U
6	12	50	1	U	U
7	15	50	1	U	U
8	20	50	1	U	U

in which $\epsilon = 0.001$.

For **CASE B** we have a fixed geometrical configuration (i.e. $a_{1,1} \equiv 1$). An overview of the stability results is shown in Table 3. There we present for different Rayleigh numbers a stability overview based on the eigenvalue approach (Figure 6 (B)) and on direct numerical simulations using the quantity $E(t)$ (Figure 7 (B)). This time the stability behavior in Table 3 is derived from Figure 6 (B). Figure 9 shows direct numerical simulations of the salt concentration for the cases 1, 7, and 8. In those simulations we have used $\Delta\mathcal{M} = 0.15$, $\Delta t = 0.005$, and $N = 30000$ which corresponds to a (dimensionless) time horizon of 150.

Figure 7 (A) shows E over time for an ℓ_x range as given in Table 2. For $\ell_x = 5\pi$ we have a monotonically decreasing E , which means that the solution remains close to the ground state $C_0(z, t)$ and the discrete $L^2(\Omega_A)$ -norm of the perturbation decays in time. Growth is observed for $\ell_x > 7\pi$: there exists a time t^* and t^{**} for which

$$\begin{aligned} E'(t) &< 0 & \text{for } 0 < t < t^* , \\ E'(t) &> 0 & \text{for } t^* < t < t^{**} , \\ E'(t) &< 0 & \text{for } t > t^{**} . \end{aligned}$$

Here $'$ denotes the derivative with respect to time. The interval (t^*, t^{**}) is referred to as the unstable episode obtained by direct numerical simulations. Numerically we find that it is consistent with

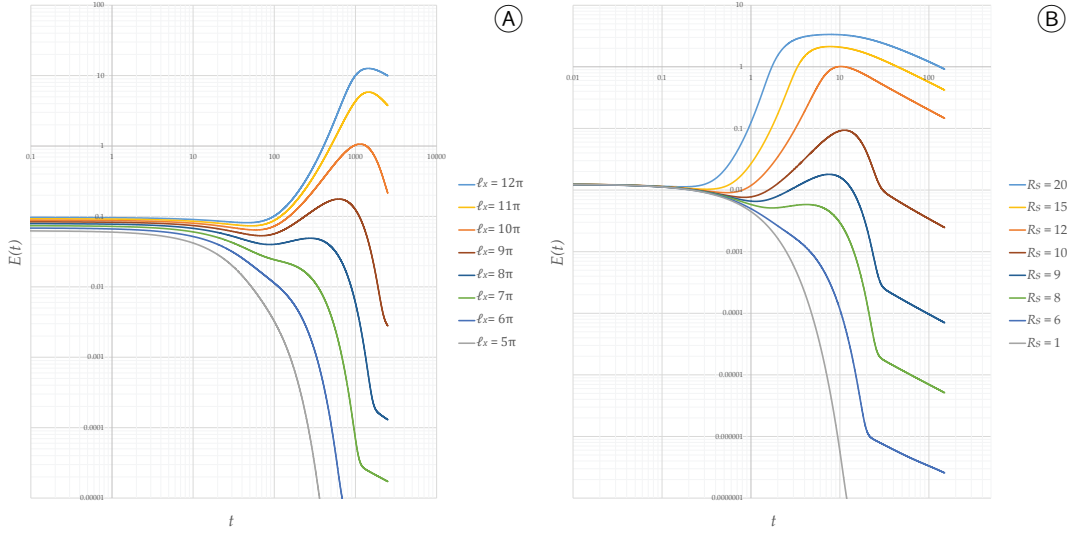


Fig. 7: $E(t)$ on a log-log scale, for a homogeneous porous column. (A): **CASE A**, with data from Table 2. (B): **CASE B**, with data from Table 3.

the theoretical predicted range $(t_L(a_{1,1}), t_R(a_{1,1}))$, as given by Figure 6:

$$t_L(a_{1,1}) < t^* < t^{**} < t_R(a_{1,1}) .$$

For example, $\ell_x = 8\pi$ ($a_{1,1} = 0.125$) yields

$$(t_L(0.125), t_R(0.125)) = (13, 290) \quad \text{and} \quad (t^*, t^{**}) = (88, 277) .$$

Note that the threshold value $\ell_x = 7\pi$ is slightly larger than the theoretically predicted value of 6.82π .

Figure 7 (B) shows again E over time for an R_S range as given by Table 3. For $R_S = 1$ we have a monotonically decreasing $E(t)$, which means that the solution remains close to the ground state $C_0(z, t)$ and the discrete $L^2(\Omega_B)$ -norm of the perturbation decays in time. For the other two numerical experiments we find

$$\begin{aligned} (t_L(1), t_R(1)) &= (0.024, 39.25) \quad \text{and} \quad (t^*, t^{**}) = (0.32, 7.7) \quad \text{for } R_S = 15, a_{1,1} = 1 , \\ (t_L(1), t_R(1)) &= (0.11, 84) \quad \text{and} \quad (t^*, t^{**}) = (0.17, 7.8) \quad \text{for } R_S = 20, a_{1,1} = 1 . \end{aligned}$$

Also note that for those two system Rayleigh numbers, the used wavenumber of the initial perturbation ($a = 1$) is relatively far away from its critical value 2.2 and 2.93, respectively (see also Figure 6, right). As a result, the numerical bounds are not sharp with respect to their corresponding theoretical bounds.

Remark 7 As can be observed in Figure 7, quantity E decays to zero for large times, indicating that the system will return to some one-dimensional profile. However, this profile is not the one-dimensional ground state solution due to the transient convective instability. The long-term one-dimensional profile can still be considered as the solution of (17), but with an adjusted initial condition.

The concentration profiles as function of time are presented in Figure 8 for **CASE A**, and Figure 9 for **CASE B**. In both cases, we observe for the selected cases episodes of (transient) instability. The instability introduces convective flow for a specific period of time. This convective dominated regime enhances the diffusive downward transport of salt into the porous column. The speed of the finger tips is higher for larger domains (**CASE A**) and larger Rayleigh numbers (**CASE B**). Along Γ_3 , near the corners, there is upward flow. Such upward flow was observed experimentally by Webster et al. (1996). The combination of the concentration boundary condition along Γ_3 and upward flow leads near the corners to a thin concentration boundary layer.

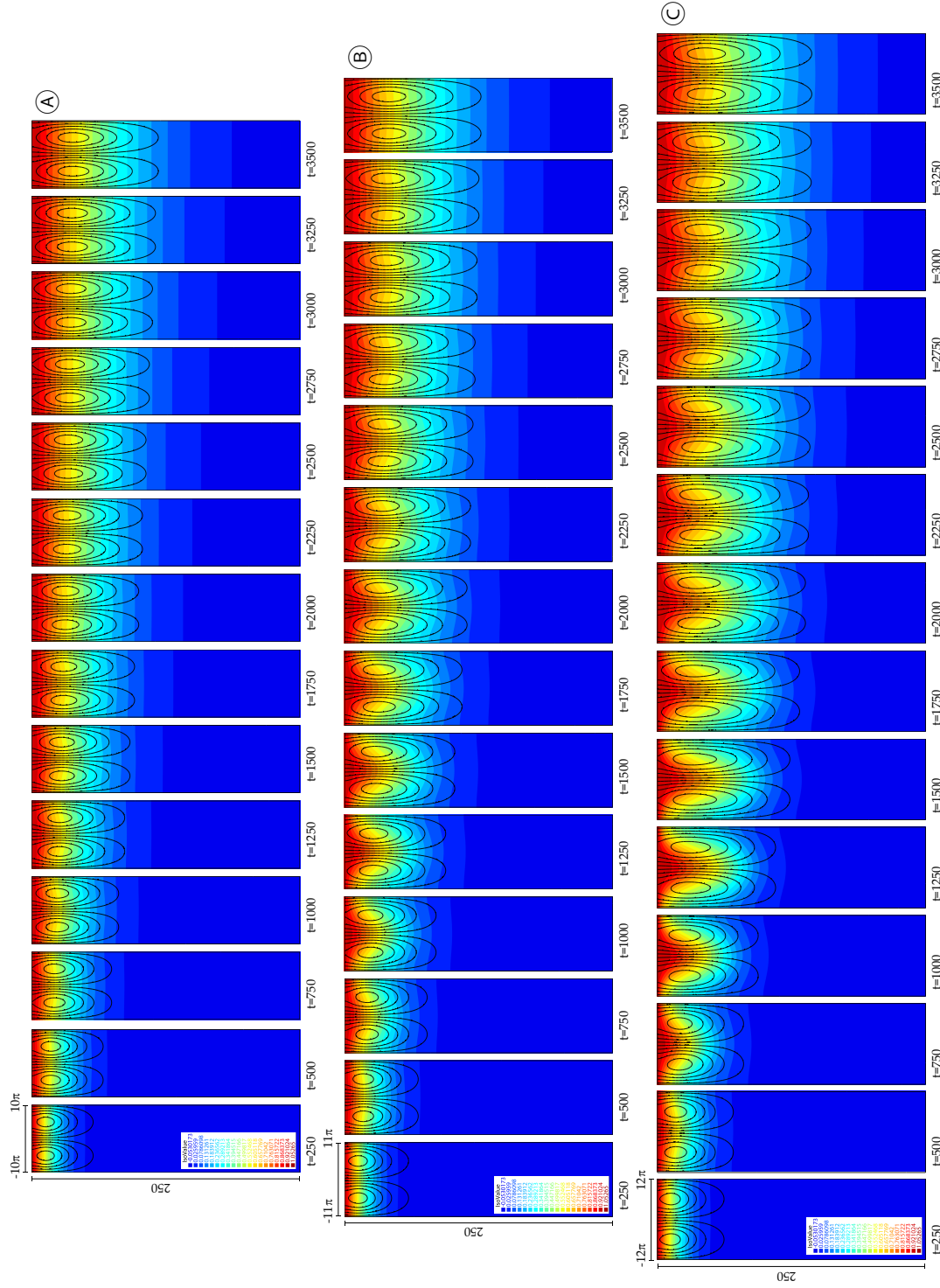


Fig. 8: Concentration profiles as function of time for **CASE A:** (A) $\ell_x = 10\pi$, (B) $\ell_x = 11\pi$ and (C) $\ell_x = 12\pi$. For those geometrical configurations, initial perturbations decay, grow and eventually decay again in time, exhibiting an unstable episode.

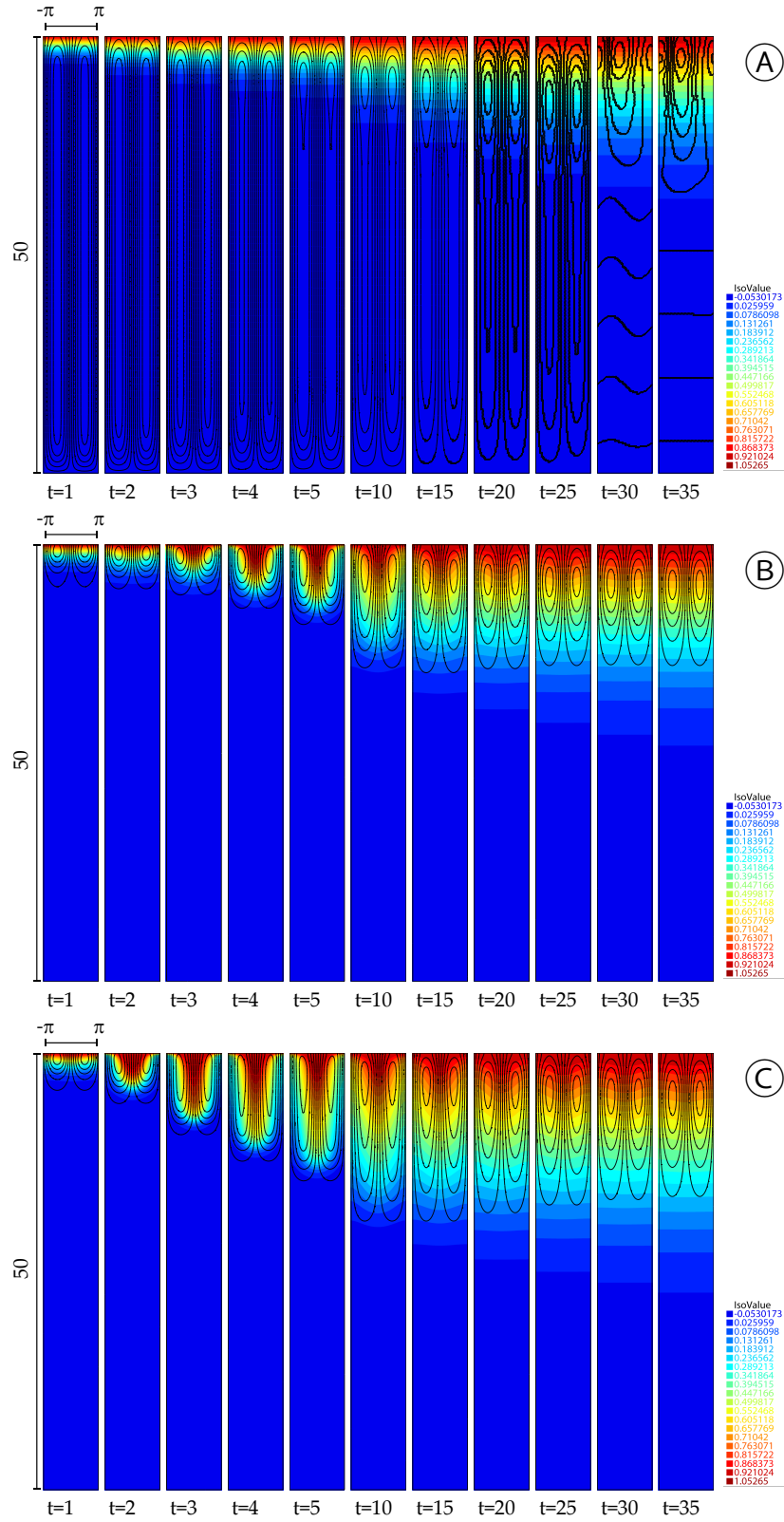


Fig. 9: Concentration profiles as function of time for **CASE B**: (A) $R_S = 1.0$. This represents a stable configuration in which only diffusive transport takes place. (B) and (C): $R_S = 15$ and $R_S = 20$ respectively. Here there exists episodes of instability.

6.2. Layered soils

Until now, we assumed for **CASE A** and **CASE B** homogeneous media only. We now consider, for **CASE B** only, a heterogeneous medium having three layers, in which one of the layers has a reduced permeability. The layered permeability structure is represented by

$$\kappa(z) = \begin{cases} \underline{\kappa} & \text{in } \{(x, z) \in \widehat{\Omega}_A^d : d_1 < z < d_2, 0 < d_1 < d_2 < d\} \\ 1 & \text{elsewhere.} \end{cases}$$

Since the dimensional permeability is scaled by $\bar{\kappa}$, we have $\underline{\kappa} \leq 1$.

In each vertical layer, the permeability is assumed to be constant. Hence, in each layer a ‘local’ system Rayleigh number, denoted by $R_S^{\underline{\kappa}}$, can be defined. We consider two values for $\underline{\kappa}$: 0.5 and 0.05. The layer parameters d_1 , d_2 , and d are given by 3, 5 and 50 respectively. The initial condition is again symmetrically perturbed by (86) with $a_{1,1} \equiv 1$.

Numerical simulations for this three-layer setup are shown in Figure 10 for a system Rayleigh number $R_S = 20$. This Rayleigh number corresponds to a system with an unstable episode, see also Table 3: The perturbation develops in layer *I* until the finger tip reaches $z = d_1$, the start of the layer *II*. The finger pattern travels through layer *II* and as can be clearly seen for $t = 4$ and $t = 5$, the finger has become more narrow. For the special case in which layer *II* is almost impermeable, the convection cells in layer *I* develop and disappear over time again, whereas in layer *III* a reverse process takes place: no convection cells exist for small times, but for larger times they appear. However, for $R_S = 20$ those convection cells in layer *III* are too weak to create a finger structure. Ultimately, the finger structure disappears and diffusion becomes the dominant process again.

The effect of $\underline{\kappa}$ is also reflected by the stability measure $E(t)$, as shown in Figure 11. For smaller values of $\underline{\kappa}$, we observe a shorter episode of instability (t^* , t^{**}). Further, the magnitude of the growing instability is, as to be expected, less pronounced because the almost impermeable layer blocks the convective transport of the finger to deeper regions.

7. INTERPRETATION IN TERMS OF DIMENSIONAL PHYSICAL PARAMETERS

Thus far the entire analytical and numerical analysis has been carried out in terms of dimensionless variables. In the concrete, dimensional world, the analysis covers a vast number of special cases. To interpret the results for specific cases, one has to return to dimensional variables by multiplying the dimensionless variables by the corresponding characteristic scales. We use subscript D to denote the dimensional variables. Specifically, multiplying the spatial dimensionless coordinates x, y, z by the characteristic length scale L_c gives spatial dimensional coordinates x_D, y_D, z_D and multiplying the dimensionless time t by the characteristic time T_c gives the dimensional time t_D :

$$\{x_D, y_D, z_D\} = \{L_c x, L_c y, L_c z\} \quad \text{and} \quad t_D = T_c t. \quad (87)$$

Also of interest are the dimensional wave number a_D and wave length λ_D given by

$$a_D = a L_c^{-1} \quad \text{and} \quad \lambda_D = L_c \lambda \quad \text{where} \quad \lambda_D = \frac{2\pi}{a_D}. \quad (88)$$

To illustrate the conversion, we use the typical values of quantities listed in Table 4. Introducing the values of ϕ , ρ_f , ρ_s , g , μ , κ_c and D from Table 4 in Table 1 gives the characteristic scales for discharge (Q_c), length (L_c), time (T_c), pressure (P_c), and Rayleigh number R . If κ_c is left to be chosen, then Table 5 results.

Motivated by solute exchange by convection within estuarine sediments, Webster et al. (1996) made an experimental and numerical analysis of unstable penetration of ponded saline water into a layer of porous medium initially saturated with fresh water, using $\kappa_c = 5.4 \times 10^{-11} \text{m}^2$.

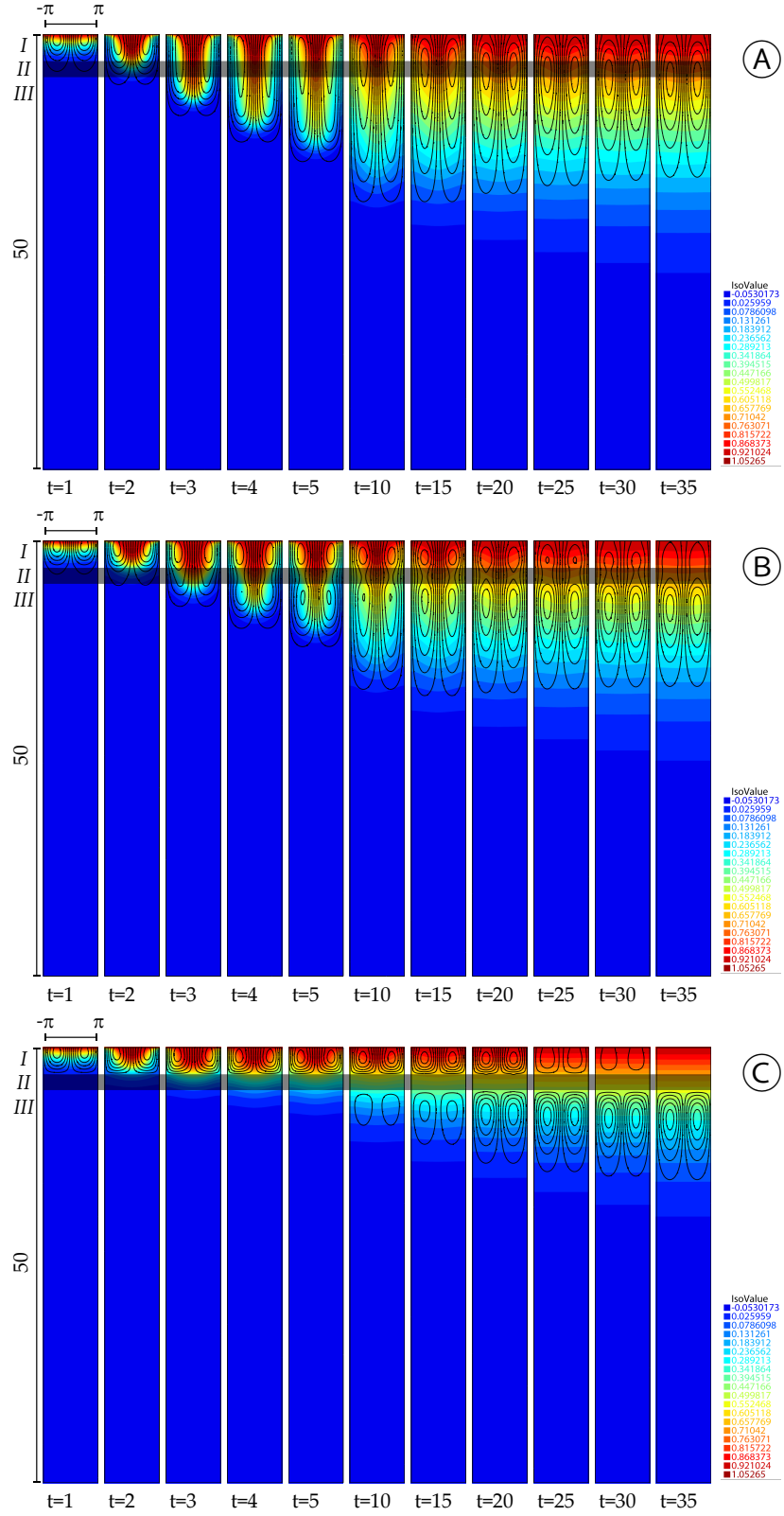


Fig. 10: Concentration profiles as function of time for **CASE B** with $R_S = 20$. The porous medium is divided in three layers: I, II, and III. Layer II has a different permeability, and is shown in gray. Layers I and II have permeability $\kappa = 1$. The following permeability values for layer II are considered: $\kappa = 1.0$ (A; for reference), $\kappa = 0.5$ (B), and $\kappa = 0.05$ (C).

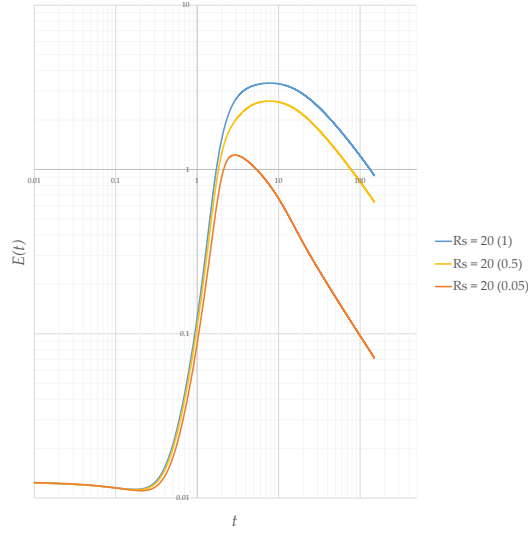


Fig. 11: $E(t)$ on a log-log scale, for a heterogeneous porous column. The numbers between brackets represent $\underline{\kappa}$ in layer II.

Post and Kooi (2003); Post (2004) numerically modeled the rates of salinization in the high-permeability Pleistocene coastal aquifer of the Netherlands, using $\kappa_c = 10^{-11} \text{m}^2$. More or less in line with those values of κ_c , we consider in the first instance a porous medium with $\kappa_c = 10^{-12} \text{m}^2$ and subsequently infer results for coarser porous media with larger values of κ_c and for finer porous media with lower values of κ_c . Substitution of $\kappa_c = 10^{-12} \text{m}^2$ in Table 5 gives the values of Q_c , L_c , T_c , P_c and R in Table 6.

Table 4: Typical values or ranges of values of physical quantities, using the International System of Units (SI).

ϕ	porosity	--	0.35
ρ_f	fresh water density	ML^{-3}	1000 kg/m ³
ρ_s	sea water density	ML^{-3}	1025 kg/m ³
g	acceleration of gravity	LT^{-2}	10 m/s
μ	viscosity	$ML^{-1}T^{-1}$	10^{-3} kg/m/s
κ_c	permeability	L^2	in range 10^{-10} – 10^{-16} m ²
D	diffusivity	L^2T^{-1}	1.5×10^{-9} m ² /s

7.1. Laterally unbounded **CASE A**

According to Table 5 (left column), for **CASE A** the Rayleigh number $R = 1$ and the characteristic scales Q_c , L_c , T_c and P_c are proportional to powers of the characteristic permeability κ_c . To the wide range of values of κ_c in Table 4 from 10^{-10} m² for very coarse porous media to 10^{-16} m² for very fine porous media correspond wide ranges of the values of the scales Q_c , L_c , T_c and P_c .

The approximate equality of the stability bounds from the linear and energy stability led for **CASE A** to the dimensionless condition (74) for L^2 -stability:

$$a_{1,1} = \sqrt{\left(\frac{\pi}{\ell_x}\right)^2 + \left(\frac{\pi}{\ell_y}\right)^2} \geq \frac{1}{6.82}. \quad (89)$$

Table 5: Overview of parameters for **CASE A** and **CASE B** obtained by substitution in Table 1 the values for ϕ , ρ_f , ρ_s , g , μ and D from Table 4.

laterally unbounded CASE A	laterally bounded CASE B
$Q_c = 250 \times 10^3 \kappa_c$	
$L_c = 6 \times 10^{-15} \kappa_c^{-1}$	$L_c = 0.3183 L_x$
$T_c = 8.4 \times 10^{-21} \kappa_c^{-2}$	$T_c = 0.02364 \times 10^9 L_x^2$
$P_c = 1.5 \times 10^{-12} \kappa_c^{-1}$	$P_c = 79.58 L_x$
$R = 1$	$R = 53.05 \times 10^{12} \kappa_c L_x$

Table 6: Overview of parameters for **CASE A** and **CASE B** obtained by setting $\kappa_c = 10^{-12} \text{m}^2$ in Table 5.

laterally unbounded CASE A	laterally bounded CASE B
$Q_c = 250 \times 10^{-9}$	
$L_c = 6 \times 10^{-3}$	$L_c = 0.3183 L_x$
$T_c = 8.4 \times 10^3$	$T_c = 0.02364 \times 10^9 L_x^2$
$P_c = 1.5$	$P_c = 79.58 L_x$
$R = 1$	$R = 53.05 L_x$

In dimensional variables, this is equivalent to:

$$a_{1,1,D} = \sqrt{\left(\frac{\pi}{\ell_x}\right)^2 + \left(\frac{\pi}{\ell_y}\right)^2} L_c^{-1} \geq \frac{L_c^{-1}}{6.82} ,$$

where L_c given by the left hand side column in Table 5. Only perturbations with wave numbers that are smaller will grow and decay in a certain time interval $t_{L,D} = T_c t_L$ to $t_{R,D} = T_c t_R$, where t_L and t_R can be read from Figure 6 (A). The corresponding condition for the wave length is derived from (89):

$$\lambda_D = \frac{2\pi}{a_{1,1,D}} \leq 42.85 L_c .$$

Only perturbations with wave lengths that are larger will grow and decay in a certain time interval $t_{L,D} = T_c t_L$ to $t_{R,D} = T_c t_R$.

Using the values in the left column of Table 6, Table 2 is translated to dimensional form by multiplying ℓ_x and d by $L_c = 6 \times 10^{-3} \text{m}$. Figure 7 (A) is translated to dimensional form by multiplying the t -axis by $T_c = 8.4 \times 10^3 \text{s}$ and ℓ_x by $L_c = 6 \times 10^{-3} \text{m}$. The dimensionless numerical results shown in Figure 8 are be translated to dimensional form by multiplying the depth $d = 250$ and the column width $\ell_x = 10, 11$, and 12 by $L_c = 6 \times 10^{-3} \text{m}$, giving $d_D = 1.5 \text{m}$ and $\ell_{xD} = 6.0, 6.6, 7.2 \text{cm}$ and the times $t = 200, 500, \dots, 3500$ by $T_c = 8.4 \times 10^3 \text{s}$ giving $t_D = 24.25, 48.50, \dots, 339$ days.

For a finer porous medium with a 10-fold smaller value $\kappa_c = 10^{-13} \text{m}^2$, the values of $\ell_{xD} = 60, 66, 72 \text{cm}$ are 10-fold larger and the values of $t_D = 6.64, 13.29, \dots, 93.01$ years are 100-fold larger.

For a coarser porous medium with a 10-fold larger value $\kappa_c = 10^{-11} \text{m}^2$, the values of $\ell_{xD} = 6.0, 6.6, 7.2 \text{mm}$ are 10-fold smaller and the values of $t_D = 0.2425, 0.4850, \dots, 3.3950$ days are 100-fold smaller. For still coarser porous media, the width of the fingers will eventually be of the same order as as the size of the pores, so that the limit of the Darcy model will be approached.

7.2. Laterally bounded CASE B

According to Table 5 (right column), for **CASE B**, the characteristic length scale L_c and pressure scale P_c and the Rayleigh number R depend linearly on the horizontal length scale L_x , while for the time scale T_c this dependence is quadratic; further, the flux scale Q_c and R are proportional to κ_c . With the wide range of values of κ_c in Table 4, the corresponding wide range of values of the Rayleigh number R is $53.05 \times 10^2 L_x$ for very coarse porous media to $53.05 \times 10^{-4} L_x$ for very fine porous media.

From the expression for R in Table 5 (right column) it follows that:

$$L_x = 1.89 \times 10^{-14} \kappa_c^{-1} R \quad \text{m} . \quad (90)$$

Introducing (90) in the expressions for characteristic length scale L_c and time scale T_c in Table 5 (right column) gives:

$$L_c = 6.02 \times 10^{-15} \kappa_c^{-1} R \quad \text{m}, \quad \text{i.e.} \quad L_c = 6.02 \times 10^{-3} R \quad \text{m} \quad \text{if} \quad \kappa_c = 10^{-12} \quad \text{m}^2, \quad (91)$$

and

$$T_c = 8.323 \times 10^{-21} \kappa_c^{-1} R^2 \quad \text{s}, \quad \text{i.e.} \quad T_c = 8.323 \times 10^{-9} R^2 \quad \text{s} \quad \text{if} \quad \kappa_c = 10^{-12} \quad \text{m}^2. \quad (92)$$

For a finer porous medium with a 10-fold smaller permeability $\kappa_c = 10^{-13} \text{m}^2$, the values of L_c are 10-fold larger and the values of T_c are 100-fold larger. For a coarser porous medium with a 10-fold larger permeability $\kappa_c = 10^{-11} \text{m}^2$, the values of L_c are 10-fold smaller and the values of T_c are 100-fold smaller.

The approximate equality of the stability bounds from the linear and energy methods led for **CASE B** to the dimensionless condition (75) for L^2 -stability:

$$a_{1,1} = \sqrt{1 + \left(\frac{1}{\ell}\right)^2} \geq \frac{R_S}{6.82}. \quad (93)$$

Here ℓ represents the aspect ratio of a rectangular box $(0, L_x) \times (0, L_y)$, i.e. $\ell = L_y/L_x$. In dimensional variables this is equivalent to:

$$a_{1,1,D} = \sqrt{1 + \left(\frac{L_x}{L_y}\right)^2} L_c^{-1} \geq \frac{R_S L_c^{-1}}{6.82},$$

where L_c given by the right hand side column in Table 5. Only perturbations with wave numbers that are smaller will grow and decay in a certain time interval $t_{L,D} = T_c t_L$ to $t_{R,D} = T_c t_R$, where t_L and t_R , for a given R_S , can be read from Figure 6 (B). The corresponding condition for the wave length is derived from (93):

$$\lambda_D = \frac{2\pi}{a_{1,1,D}} \leq 42.85 L_c R_S^{-1}.$$

Only perturbations with wave lengths that are larger will grow and decay in a certain time interval $t_{L,D} = T_c t_L$ to $t_{R,D} = T_c t_R$.

Criterion (93) means that in a rectangular column the ground state is L^2 -stable provided

$$R_S < 6.82 \sqrt{1 + \frac{1}{\ell^2}}, \quad (94)$$

or, using Table 1,

$$\frac{\kappa_c (\rho_s - \rho_f) g L_x}{\mu D} < 21.43 \sqrt{1 + \frac{1}{\ell^2}}. \quad (95)$$

When the cross-section is a square, i.e. $\ell = 1$, L^2 -stability holds if

$$\frac{\kappa_c(\rho_s - \rho_f)gL_x}{\mu D} < 30.30 . \quad (96)$$

In practice, experiments are often performed in cylindrical columns see e.g. Wooding (1959). Suppose we consider in **CASE B** a semi-infinite cylinder of radius \mathcal{R} . This radius will then be used, conform **CASE B**, in the characteristic length scale $L_c = \frac{\mathcal{R}}{\pi}$. Then instead of (36), we have

$$\Omega_B = \{(x, y, z) : \sqrt{x^2 + y^2} < \pi, z > 0\} .$$

As in (33), we need to consider now the eigenvalue problem

$$\begin{cases} \Delta_{\perp} f + a^2 f = 0 & \text{in } \{(r, \varphi) : 0 < r < \pi, 0 \leq \varphi < 2\pi\} , \\ \nabla_{\perp} f \cdot \mathbf{n}|_{r=\pi} = 0 . \end{cases} \quad (97a)$$

$$(97b)$$

In particular, for the given geometry, one needs to determine the smallest $a > 0$ for which the eigenvalue problem (97) has a solution. As explained in Section 6.1.2, this solution describes the most unstable mode.

Solutions of (97a) are found by separation of variables:

$$f(r, \varphi) = J_m(ar) \cos m\varphi , \quad (98)$$

where J_m denotes the Bessel function of the first kind of order m . Boundary condition (97b) is satisfied when

$$J'_m(a\pi) = 0 \quad (' \text{ denotes differentiation}) ,$$

or

$$a = a_{n,m} = \frac{\beta_{n,m}}{\pi} \quad n = 1, 2, \dots \quad \text{and} \quad m = 0, 1, 2, \dots ,$$

where $\beta_{n,m}$ is the n^{th} positive zero of J'_m . Clearly

$$\min_{\substack{n=1,2,\dots \\ m=0,1,2,\dots}} \beta_{n,m} = \min_{m=0,1,2,\dots} \beta_{1,m} .$$

Looking up the zeros of J'_m in Weisstein gives

$$\min_{m=0,1,2,\dots} \beta_{1,m} = \beta_{1,1} = 1.8412 .$$

This yields $a = \hat{a}_{1,1} = \frac{\beta_{1,1}}{\pi}$ as the smallest $a > 0$ for which (97) has a non-trivial solution. Using $\hat{a}_{1,1}$ in (93) yields for cylinders the stability criterion

$$\frac{\kappa_c(\rho_s - \rho_f)g\mathcal{R}}{\mu D} < 12.56 . \quad (99)$$

Flow patterns corresponding to (98) are discussed in Appendix B.

The dimensionless time axis of Figure 7 for the stability measures E is translated to a dimensional time axis by multiplying it by the characteristic time T_c given by (92). Since T_c is proportional to R_S^2 , the tendency is that the transient instability occurs later for larger R_S .

The dimensionless numerical results shown in Figure 9 are be translated to dimensional form by multiplying the dimensionless depth $d = 75$ and column width 2π by the linearly R dependent characteristic length L_c given by (91) and multiplying the dimensionless times $t = 1, 2, 3, 4, 5, 10, 15, 20, 30, 35$ by the quadratically R dependent characteristic time T_c is given by (92).

8. CONCLUSIONS

We analyzed the stability of the density-depth profiles in porous media initially saturated with a stagnant fluid of density ρ_f after flooding with a salt solution of higher density ρ_s . Such flooding of coastal areas with seawater regularly occurs, either due to climatological circumstances or as part of engineering projects. Historically, flooding was sometimes also imposed as a defensive measure.

We used the standard convection-diffusion equation subject to the so-called Boussinesq approximation. The variable density arising from the diffusive penetration of salt then affects the fluid flow only by its effect on the gravitational driving force in the Darcy equation. To explore the density difference induced flow, we used both analytical and numerical methods. We assumed the depth of the porous medium to be infinite in the analytical approaches and finite in the numerical analyses. Geometrically, two cases were distinguished: the laterally unbounded **CASE A** and the laterally bounded **CASE B**. The ratio of the diffusivity D and the density difference induced gravitational shear flow $(\kappa_c/\mu)(\rho_s - \rho_f)g$ is an intrinsic length scale of the problem. In the unbounded **CASE A**, this is the only length scale and using it leads to a dimensionless formulation with Rayleigh number $R=1$. In the bounded **CASE B**, the lateral geometry provides another length scale and using it leads to a dimensionless formulation with a Rayleigh number R given by the ratio of the geometric and intrinsic length scales.

The analytical approaches lead to general statements about stability or instability under given circumstances. The background diffusion of salt from the ponded layer into the fluid saturated porous medium, the so-called ground state solution, is expressed in terms of the Boltzmann similarity variable z/\sqrt{t} . Three methods were used to study the stability of this ground state, the first two based on the linearized perturbation equation for the concentration and the third based on the full nonlinear perturbation equation. They yield eigenvalue problems that contain the horizontal wave number a and the Rayleigh number R as parameters. For homogeneous soils, using the Boltzmann variable z/\sqrt{t} , the time t can be eliminated from the eigenvalue problem by absorbing \sqrt{t} into a and R : $b = a\sqrt{t}$ and $R^* = R\sqrt{t}$. This time-independence of the eigenvalue problems leads to the three stability curves relating the transformed Rayleigh number R^* and wave number b shown in Figures 4.

For the first linear approach, the surface spatial density gradient was used as an approximation of the entire background density profile. This results in a crude estimate of the L^2 -norm of the concentration shown in Figure 3, with the perturbation at first growing, but eventually decaying in time.

For the second linear approach, the full ground state solution subject to the restriction that the ground state slowly evolves in time was used: the so-called ‘frozen profile approximation’. This method of linearized stability only holds for infinitesimal perturbations. To get results for larger perturbations, a specific version of the variational energy method was used to estimate from the full nonlinear perturbation equation the L^2 -norm of the concentration perturbation c . The stability curve $R_L^*(b)$ resulting from the linearized perturbation approach is surprisingly close to the stability curve $R_E^*(b)$ resulting from the variational energy method. This means that there is only a small gap between stability curves $R_L^*(b)$ and $R_E^*(b)$ in which small perturbations vanish, but large perturbations may grow. In this respect the current problem differs markedly from the problem with an imposed vertical background flow analyzed earlier (van Duijn et al., 2002), for which the gap was found to be large. The ponded problem considered in this paper has a sharp stability bound: below R_E^* there is unconditional stability and above R_L^* instability. The curve $R_M^* = b^2$ resulting from the first linear approach is shown to be a lower bound for both R_L^* and R_E^* .

For both **CASE A** and **CASE B**, the \sqrt{t} transformed system Rayleigh number $R_S^* = R_S \sqrt{t}$ is a linear function of the \sqrt{t} transformed system wave number b . To determine the stability for a particular wave number b - system Rayleigh number R_S combination, three cases can be distinguished, depending on whether the straight line $R_S^* = ab$ lies entirely below R_E^* , touches R_E^* at $b = b^*$, or

intersects R_E^* at b_L^* and b_R^* . Numerically it was found that the line with $\alpha = \alpha_{\text{crit}} = 3.85$ touches R_E^* at $(R^*, b) = (3.89, 1.01)$. For $\alpha < \alpha_{\text{crit}}$ the ground state is unconditionally stable. For $\alpha > \alpha_{\text{crit}}$ the ground state is unconditionally stable for $b < b_L^*$ and $b > b_R^*$ and experiences an episode of transient instability for $b_L^* < b < b_R^*$.

In Sub-subsection 6.1.2, by reversal of the Boltzmann transformations the analytical stability results were interpreted for both **CASE A** and **CASE B** in terms of system Rayleigh number R_S , wave number a and time t (see in particular Figure 6). In Sub-subsection 6.1.3 these results were compared with direct numerical simulations based on the full, spatially 2-D flow and transport equations in x, z, t , with a range of lateral length scales for **CASE A** (see Table 2) and a range of system Rayleigh numbers for **CASE B** (see Table 3). Qualitatively, the numerical results turned out to agree with the eigenvalue results. They also clearly show the details of the episodic instabilities as function of the wave number for **CASE A** and of the wave number and system Rayleigh number for **CASE B**. The episodic instability clearly leads to deeper penetration of salt.

The case of a layered soil was considered in Section 6.2. There the Boltzmann transformation does not apply and one has to rely on numerical techniques. Results of the salt distribution in an aquifer having a horizontal and lower permeable layer are shown in Figure 10. This figure clearly demonstrates the effect of the layer acting as a barrier for the flow and thus for the salt transport. This illustrates the protective role of low-permeability layers against invasion of saline water into underlying fresh water mentioned in the Introduction.

In Section 7 it was shown how the dimensionless analytical and numerical results can be translated to dimensional form and thereby interpreted in terms of dimensional length and time scales. This provided an opportunity to highlight the important role of the permeability of the porous medium: the higher the permeability the stronger is the destabilizing influence of the gravitational force.

The laterally bounded **CASE B** with its no-flow and no-solute transport conditions at the vertical boundary Γ_v is not restricted to rectangles and hexagons, but covers columns of any shape, such as the cylinders discussed at the end of Section 7. Thus it provides a good basis for the interpretation of experiments in columns. In particular, column experiments are stable provided, see also (95), (96), and (99):

- (i) Rectangular cross-section $(0, L_x) \times (0, L_y)$

$$\frac{\kappa_c(\rho_s - \rho_f)gL_x}{\mu D} < 21.43 \sqrt{1 + \frac{1}{\ell^2}}, \quad \ell = \frac{L_y}{L_x};$$

- (ii) Square cross-section $(0, L_x) \times (0, L_x)$

$$\frac{\kappa_c(\rho_s - \rho_f)gL_x}{\mu D} < 30.30;$$

- (iii) Circular cross-section with radius \mathcal{R}

$$\frac{\kappa_c(\rho_s - \rho_f)g\mathcal{R}}{\mu D} < 12.56.$$

A. CELL PATTERNS (CHANDRASEKHAR (1981))

Though a detailed study of the periodic horizontal structure is outside the scope of this paper, some observations seem useful as background material.

By the nature of the ponded-induced flow, there is no preferential direction in the horizontal (x, y) -plane. Hence it is 'reasonable' to argue that in the marginal state, the horizontal plane is divided or tessellated into regular polygons $\mathcal{T} \subset \mathbb{R}^2$ so that the vertical walls of the cells

$$\Omega_A = \{(x, y, z) : (x, y) \in \mathcal{T}, z > 0\} \quad (100)$$

are surfaces where periodic conditions hold for the concentration, discharge and pressure.

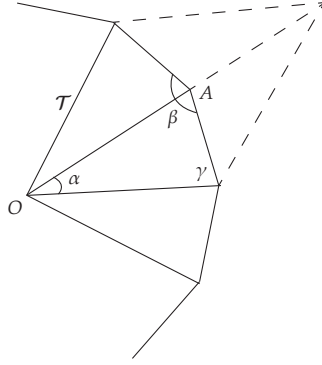


Fig. 12: Regular k -polygon.

What are the possibilities? Suppose \mathcal{T} is a regular k -polygon as in Figure 12. Then clearly $\alpha = \frac{2\pi}{k}$ and $\beta = \gamma = \frac{1}{2}(\pi - \frac{2\pi}{k}) = \frac{\pi}{2}(1 - \frac{2}{k})$. Hence the angle at vertex A is given by $\pi(1 - \frac{2}{k})$. Now suppose that l -polygon vertices meet at A . Then

$$\pi(1 - \frac{2}{k}) = \frac{2\pi}{l}$$

or

$$\frac{1}{k} + \frac{1}{l} = \frac{1}{2} \quad \text{with } k, l \in \mathbb{N}.$$

This implies three possibilities

$k = 3, l = 6$	regular triangle
$k = 4, l = 4$	rectangle or square
$k = 6, l = 3$	regular hexagon

Later we show that a hexagonal pattern contains a periodic pattern of triangles. Hence it suffices to consider rectangles and hexagons only.

What kind of flow patterns fit in these structures? In the linear approach, having marginal instabilities, we have for c , \mathbf{u} and p the equations (39). We seek solutions of this linear system of the form

$$\{c, \mathbf{u}, p\}(x, y, z, t) = \{c, \mathbf{u}, p\}(z, t)f(x, y), \quad (101)$$

where f describes the horizontal periodic behaviour. It satisfies

$$\begin{cases} \Delta_{\perp} f + a^2 f = 0 & \text{in } \mathcal{T} \text{ (or } \mathbb{R}^2 \text{)}, \\ f \text{ is } \mathcal{T}\text{-periodic,} \end{cases} \quad (102)$$

where a is a given parameter. Using (102) in (39) gives

$$\left\{ \begin{array}{l} \frac{\partial c}{\partial t} + R w \frac{\partial C_0}{\partial z} = \frac{\partial^2 c}{\partial z^2} - a^2 c, \end{array} \right. \quad (103)$$

$$\left\{ \begin{array}{l} -\frac{\partial}{\partial z} \left(\frac{1}{\kappa(z)} \frac{\partial w}{\partial z} \right) + \frac{1}{\kappa(z)} a^2 w = a^2 c, \end{array} \right. \quad (104)$$

$$\left\{ \begin{array}{l} \frac{\partial w}{\partial z} = -a^2 \kappa(z) p. \end{array} \right. \quad (105)$$

Thus $p(x, y, z, t) = p(z, t) f(x, y)$ with

$$p(z, t) = -\frac{1}{a^2 \kappa(z)} \frac{\partial w}{\partial z}(z, t). \quad (106)$$

The horizontal component of the Darcy equation (25b) gives

$$\frac{1}{\kappa(z)} u(x, y, z, t) = -p(z, t) \frac{\partial f}{\partial x}$$

and

$$\frac{1}{\kappa(z)} v(x, y, z, t) = -p(z, t) \frac{\partial f}{\partial y}$$

Substituting (106) results in

$$u(x, y, z, t) = u(z, t) \frac{\partial f}{\partial x}, \quad \text{with } u(z, t) = \frac{1}{a^2} \frac{\partial w}{\partial z}(z, t), \quad (107)$$

$$v(x, y, z, t) = v(z, t) \frac{\partial f}{\partial y}, \quad \text{with } v(z, t) = \frac{1}{a^2} \frac{\partial w}{\partial z}(z, t). \quad (108)$$

Thus the periodic boundary conditions for f on $\partial\mathcal{T}$ imply that the discharge field $\mathbf{u} = (u, v, w)$ is periodic with respect to \mathcal{T} .

What form for f do we choose? We first consider rectangles and next hexagons.

A.1. Rectangular tessellation

Then

$$\mathcal{T} = (-L_x, L_x) \times (-L_y, L_y) \quad (L_x, L_y > 0).$$

Now the \mathcal{T} -periodic solution of (102) is given by

$$f(x, y) = \cos a_x x \cos a_y y,$$

where $a = \sqrt{a_x^2 + a_y^2}$ and $a_x = \frac{n\pi}{L_x}$ and $a_y = \frac{m\pi}{L_y}$. Hence for the components of $\mathbf{u} = (u, v, w)$ we have

$$\left. \begin{array}{l} u(x, y, z, t) = -\frac{a_x}{a^2} \frac{\partial w}{\partial z}(z, t) \sin a_x x \cos a_y y \\ v(x, y, z, t) = -\frac{a_y}{a^2} \frac{\partial w}{\partial z}(z, t) \cos a_x x \sin a_y y \\ w(x, y, z, t) = w(z, t) \cos a_x x \cos a_y y \end{array} \right\} \quad (109)$$

where $w(z, t)$ is the solution of (39a) and (39b) under the appropriate boundary conditions. An impression of the flow patterns is shown in Figure 13.

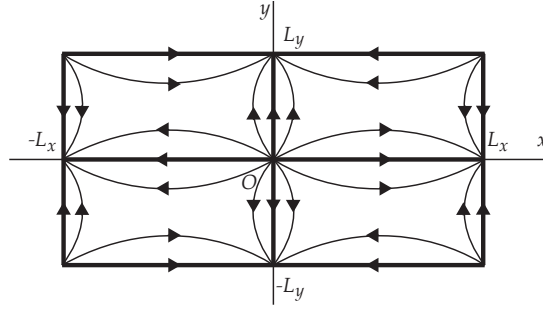


Fig. 13: Scaled horizontal flow in rectangular periodic cell: $\frac{u(x, y, z, t)}{\frac{-a_x}{a^2} \frac{\partial w}{\partial z}} = \sin a_x x \cos a_y y, \frac{v(x, y, z, t)}{\frac{-a_y}{a^2} \frac{\partial w}{\partial z}} = \cos a_x x \sin a_y y$ where $a_x = \frac{\pi}{L_x}$ and $a_y = \frac{\pi}{L_y}$.

A.2. Hexagonal tessellation

Let \mathcal{T} be the regular hexagon of which the sides have length L . An example is sketched in Figure 14. In this figure we also show the construction of a corresponding regular triangular cell of which the sides have length $\sqrt{3}L$. Inspection of the resulting discharge field shows that it suffices to consider the hexagons only.

Problem (102) was originally solved by Christopherson (1940). The solution f is given by

$$f(x, y) = \cos \frac{4\pi}{3L} \left(\frac{1}{2} \sqrt{3}x + \frac{1}{2}y \right) + \cos \frac{4\pi}{3L} \left(\frac{1}{2} \sqrt{3}x - \frac{1}{2}y \right) + \cos \frac{4\pi}{3L} y, \quad \text{and} \quad a = \frac{4\pi}{3L}. \quad (110)$$

Let us consider the corresponding discharge pattern and the periodicity of the boundary conditions. We first check the periodicity of f in (110). In polar coordinates ($x = r \cos \varphi$, $y = r \sin \varphi$) it reads

$$f(x, y) = \cos \frac{4\pi}{3L} r \sin \left(\varphi + \frac{\pi}{3} \right) + \cos \frac{4\pi}{3L} r \sin \left(\varphi + \frac{2\pi}{3} \right) + \cos \frac{4\pi}{3L} r \sin \varphi.$$

Hence

$$f(r, \varphi) = f(r, \varphi + \frac{\pi}{3}), \quad (111)$$

i.e. f is $\frac{\pi}{3}$ -rotation periodic. Further,

$$f(x + L_x^*, y) = f(x, y) \quad \text{with } L_x^* = \sqrt{3}L$$

and

$$f(x, y + L_y^*) = f(x, y) \quad \text{with } L_y^* = \sqrt{3}L.$$

In view of (111) we only need to investigate f in the triangle OMN , see Figure 14. Along the vertical boundary MN we have

$$x = \frac{1}{2} \sqrt{3}L, \quad -\frac{1}{2}L < y < \frac{1}{2}L$$

and thus

$$\begin{aligned} f|_{MN} &= \cos \left(\pi + \frac{2\pi}{3L} y \right) + \cos \left(\pi - \frac{2\pi}{3L} y \right) + \cos \frac{4\pi}{3L} y \\ &= -2 \cos \frac{2\pi}{3L} y + \cos \frac{4\pi}{3L} y \\ &= f|_{M'N'}. \end{aligned} \quad (112)$$

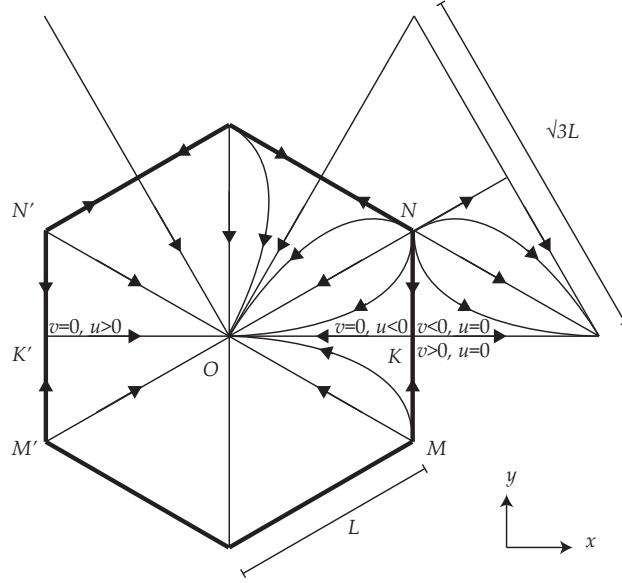


Fig. 14: Hexagonal cell in the horizontal plane. Note that a tessellation of triangular cells is covered by this case. Horizontal flow is scaled as in Figure 13.

Note that $f|_{M'N'}$ is symmetric with respect to the point K.

To compute the scaled horizontal flow components $\frac{u}{\frac{-a_x}{a^2} \frac{\partial w}{\partial z}}$ and $\frac{v}{\frac{-a_y}{a^2} \frac{\partial w}{\partial z}}$ we need to determine

$$\nabla_{\perp} f = \begin{pmatrix} \frac{\partial f}{\partial x} \\ \frac{\partial f}{\partial y} \end{pmatrix} = \begin{pmatrix} \frac{-2\pi}{\sqrt{3}L} \left\{ \sin \frac{4\pi}{3L} \left(\frac{1}{2} \sqrt{3}x + \frac{1}{2}y \right) + \sin \frac{4\pi}{3L} \left(\frac{1}{2} \sqrt{3}x - \frac{1}{2}y \right) \right\} \\ \frac{-2\pi}{3L} \left\{ \sin \frac{4\pi}{3L} \left(\frac{1}{2} \sqrt{3}x + \frac{1}{2}y \right) - \sin \frac{4\pi}{3L} \left(\frac{1}{2} \sqrt{3}x - \frac{1}{2}y \right) + 2 \sin \frac{4\pi}{3L} y \right\} \end{pmatrix}.$$

Equivalently one can write

$$\nabla_{\perp} f = \begin{pmatrix} \frac{-4\pi}{\sqrt{3}L} \sin \frac{2\pi}{\sqrt{3}L} x \cos \frac{2\pi}{3L} y \\ \frac{-4\pi}{3L} \cos \frac{2\pi}{\sqrt{3}L} x \sin \frac{2\pi}{3L} y - \sin \frac{4\pi}{3L} y \end{pmatrix}.$$

Segment MN: We have

$$\begin{aligned} \frac{\partial f}{\partial x} &= 0 \\ \frac{\partial f}{\partial y} &= \frac{4\pi}{3L} \left(\sin \frac{2\pi}{3L} y - \sin \frac{4\pi}{3L} y \right) \\ &= \frac{4\pi}{3L} \sin \frac{2\pi}{3L} y \left(1 - 2 \cos \frac{2\pi}{3L} y \right) \begin{cases} < 0 & \text{on } KN \\ > 0 & \text{on } MK \end{cases}. \end{aligned}$$

By the rotation symmetry $\nabla_{\perp} f \cdot \mathbf{n} = 0$ on $\partial\mathcal{T}$. Combined with (112) this implies that f satisfies periodic boundary conditions on $\partial\mathcal{T}$.

Segment OK: We have

$$\begin{aligned}\frac{\partial f}{\partial x} &= -\frac{4\pi}{3L} \sin \frac{2\pi}{\sqrt{3}L} x < 0 \\ \frac{\partial f}{\partial y} &= 0\end{aligned}$$

Segment ON: Then $y = x/\sqrt{3}$, $\mathbf{n} = (-1, \sqrt{3})^T$ (normal direction), $\mathbf{t} = (\sqrt{3}, 1)^T$ (tangential direction). This implies, with $\alpha = \frac{2\pi}{\sqrt{3}L}x$,

$$\nabla f|_{ON} = \begin{pmatrix} \frac{-4\pi}{\sqrt{3}L} \sin \alpha \cos \frac{1}{3}\alpha \\ \frac{-4\pi}{3L} (\cos \alpha \sin \frac{1}{3}\alpha + \sin \frac{2}{3}\alpha) \end{pmatrix},$$

where $0 < \alpha < \pi$. Consequently,

$$\frac{\partial f}{\partial n}|_{ON} = \nabla_{\perp} f \cdot \mathbf{n}|_{ON} = \frac{4\pi}{\sqrt{3}L} \left(\sin \alpha \cos \frac{1}{3}\alpha - \cos \alpha \sin \frac{1}{3}\alpha - \sin \frac{2}{3}\alpha \right) = 0$$

and

$$\frac{\partial f}{\partial t}|_{ON} = \nabla_{\perp} f \cdot \mathbf{t}|_{ON} = \frac{-4\pi}{\sqrt{3}L} \left(3 \sin \alpha \cos \frac{1}{3}\alpha + \cos \alpha \sin \frac{1}{3}\alpha + \sin \frac{2}{3}\alpha \right) < 0$$

since

$$\begin{aligned}3 \sin \alpha \cos \frac{1}{3}\alpha + \cos \alpha \sin \frac{1}{3}\alpha + \sin \frac{2}{3}\alpha &= 2 \sin \alpha \cos \frac{1}{3}\alpha + \sin \frac{4}{3}\alpha + \sin \frac{2}{3}\alpha \\ &> \sin \frac{4}{3}\alpha + \sin \frac{2}{3}\alpha = \sin \frac{2}{3}\alpha \left(2 \cos \frac{2}{3}\alpha + 1 \right),\end{aligned}$$

and $\cos \frac{2}{3}\alpha > -\frac{1}{2}$ for $0 < \alpha < \pi$. Hence the scaled flow field behaves as sketched in Figure 14.

B. LATERALLY BOUNDED, CYLINDRICAL COLUMN (WOODING (1959))

Then

$$\Omega = \{(r, \varphi) : 0 < r < \pi, 0 \leq \varphi < 2\pi\}$$

denotes the dimensionless cross-section. The horizontal periodicity is described by the eigenvalue problem

$$\begin{aligned}\Delta_{\perp} f + a^2 f &= 0 \quad \text{in } \Omega, \\ \nabla_{\perp} f \cdot \mathbf{n}|_{r=\pi} &= 0.\end{aligned}$$

As explained in Section 6 of the main text, the smallest positive a for which this problem has a nontrivial solution corresponds to the most unstable mode.

By separation of variables, see for instance Strauss (2007), one finds the eigenvalues

$$a = a_{n,m} = \frac{\beta_{n,m}}{\pi}, \quad \beta_{n,m} \text{ is the } n^{\text{th}} \text{ positive zero of } J'_m,$$

for $n = 1, 2, \dots$ and $m = 0, 1, \dots$, and the eigenfunctions

$$f_{n,m}(r, \varphi) = J_m(a_{n,m}r) \cos m\varphi,$$

where J_m is the Bessel function of the first kind and order m .

We are interested in the smallest $a_{n,m}$. By the ordering of the extrema one has

$$a_{1,m} < a_{2,m} < \cdots \quad \text{for } m=0, 1, \cdots .$$

Further, see Weisstein, the smallest values $a_{1,m}$ are ordered as

$$a_{1,1} < a_{1,2} < a_{1,0} < a_{1,3} < \cdots ,$$

with $a_{1,1} = \frac{1.8412}{\pi}$.

The horizontal flow pattern follows from

$$\mathbf{u}_{\perp,n,m}(r, \varphi, z, t) = \frac{1}{a^2} \frac{\partial w}{\partial z}(z, t) \nabla_{\perp} f_{n,m}(r, \varphi) ,$$

where

$$\begin{aligned} \nabla_{\perp} f_{n,m} &= \frac{\partial f_{n,m}}{\partial r} \mathbf{e}_r + \frac{1}{r} \frac{\partial f_{n,m}}{\partial \varphi} \mathbf{e}_{\varphi} \\ &= a_{n,m} J'_m(a_{n,m} r) \cos m\varphi \mathbf{e}_r - \frac{m}{r} J_m(a_{n,m} r) \sin m\varphi \mathbf{e}_{\varphi} . \end{aligned}$$

For the first four eigenvalues, the patterns and level sets of $f_{n,m}$ are sketched in Figure 15. The ordering of the stability and the flow patterns correspond qualitatively with the experimental results of Wooding (1959), Figure 1 and Tables 1 and 2.

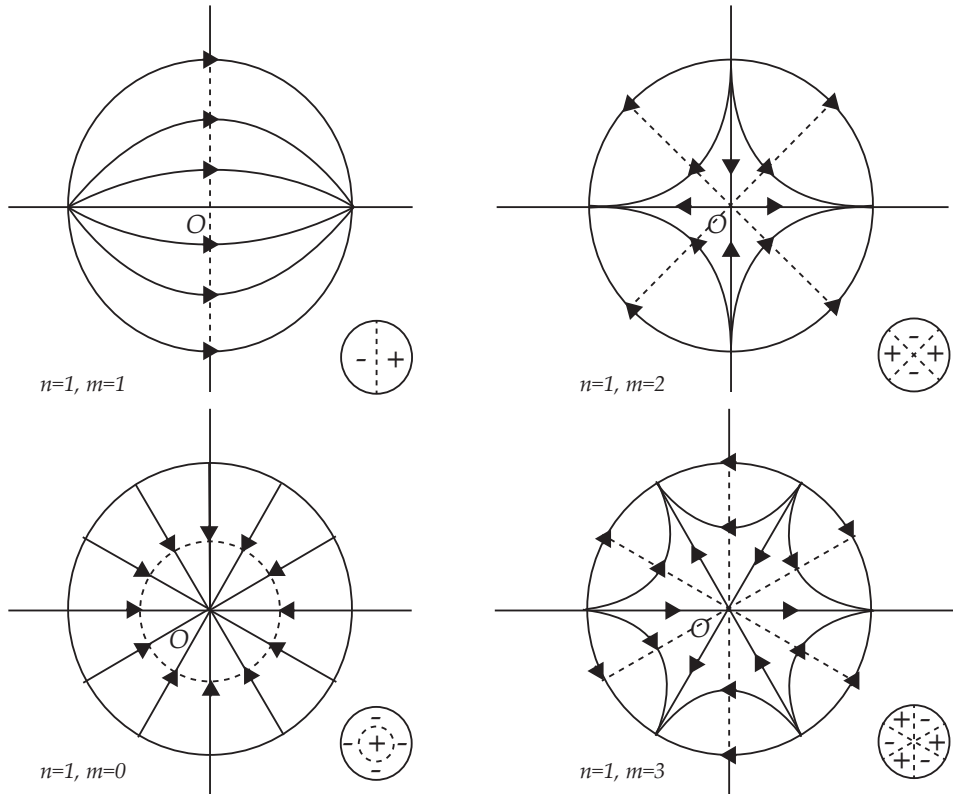


Fig. 15: Sketch of flow patterns and level set of $f_{1,m}$, for $m=1, 2, 0, 3$.

Since $w(r, \varphi, z, t) = w(z, t) f(r, \varphi)$, one visualizes downward flow in the regions where $f > 0$, and upward flow when $f < 0$. A rough indication about the ordering of stability is obtained by

computing the ratio

$$r_{n,m} = \frac{\text{length}}{\text{area}} \Big|_{f_{n,m} > 0}.$$

This gives, see also Figure 15,

$$\begin{aligned} r_{1,1} &= \frac{2\pi}{\frac{1}{2}\pi^3} = \frac{4}{\pi^2} \\ r_{1,2} &= \frac{4\pi}{\frac{1}{2}\pi^3} = 2\frac{4}{\pi^2} \\ r_{1,0} &= \frac{2\pi r_0}{\pi r_0^2} = \frac{\pi^2}{2r_0} \frac{4}{\pi^2} \approx 2.5 \frac{4}{\pi^2} \\ r_{1,3} &= \frac{6\pi}{\frac{1}{2}\pi^3} = 3\frac{4}{\pi^2} \end{aligned}$$

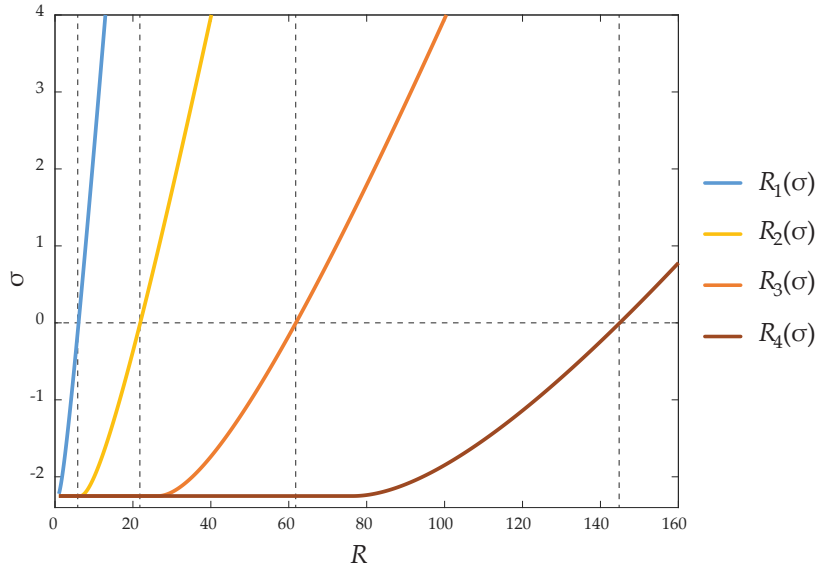


Fig. 16: Calculated eigenvalues $R_i(\sigma) := R_i(a, t, \sigma)$, $a = 1.5$, $t = 1$, for $i = 1, 2, 3, 4$.

C. EXCHANGE OF STABILITIES

Eigenvalue problem (57) is solved for a fixed wave number a and time t , and for the range $R \in [1, 20]$. The eigenvalue is discretized by means of a Chebyshev–Petrov–Galerkin method, see Appendix H. The first four eigenvalues are shown in Figure 16. Based on numerical calculations, it follows that the spectra reveal real eigenvalues only, and that $\sigma \leq 0$ for $R \leq R_i(\sigma)$ for at least $i = 1, \dots, 4$.

D. DERIVATION OF THE EULER-LAGRANGE EQUATIONS

We need to find (c, w, π) and values of R where the first variation of (65) vanishes. Let $(\delta, \epsilon) \in \mathcal{H}$ and let μ be Ω_i -periodic with

$$\frac{\partial \mu}{\partial z} = 0, \quad \text{at } z=0, \text{ for } t>0, \quad (113a)$$

$$\mu \rightarrow 0, \quad \text{as } z \rightarrow 0, \text{ for } t>0. \quad (113b)$$

Then consider

$$\begin{aligned} J(c + \delta, w + \epsilon, \pi + \mu; R) &= \int_{\Omega_i} |\nabla(c + \delta)|^2 + R \int_{\Omega_i} (c + \delta)(w + \epsilon) \frac{\partial C_0}{\partial z} + \\ &+ \int_{\Omega_i} (\pi + \mu) \left\{ \nabla \cdot \left(\frac{1}{\kappa} \nabla(w + \epsilon) \right) - \Delta_{\perp}(c + \delta) \right\} = J(c, w, \pi; R) + 2 \int_{\Omega_i} \nabla c \cdot \nabla \delta + \int_{\Omega_i} |\nabla \delta|^2 + \\ &+ R \int_{\Omega_i} \delta w \frac{\partial C_0}{\partial z} + R \int_{\Omega_i} \epsilon c \frac{\partial C_0}{\partial z} + R \int_{\Omega_i} \delta \epsilon \frac{\partial C_0}{\partial z} + \\ &+ \int_{\Omega_i} \mu \left\{ \nabla \cdot \left(\frac{1}{\kappa} \nabla w \right) - \Delta_{\perp} c \right\} + \int_{\Omega_i} \pi \left\{ \nabla \cdot \left(\frac{1}{\kappa} \nabla \epsilon \right) - \Delta_{\perp} \delta \right\} + \int_{\Omega_i} \mu \left\{ \nabla \cdot \left(\frac{1}{\kappa} \nabla \epsilon \right) - \Delta_{\perp} \delta \right\}. \end{aligned}$$

Terms that are linear in δ, ϵ , and μ form the first variation of (65) at (c, w, π) and R . Collecting these terms yields

$$2 \int_{\Omega_i} \nabla c \cdot \nabla \delta + R \int_{\Omega_i} \delta w \frac{\partial C_0}{\partial z} - \int_{\Omega_i} \pi \Delta_{\perp} \delta = 0 \quad (114a)$$

$$R \int_{\Omega_i} \epsilon c \frac{\partial C_0}{\partial z} + \int_{\Omega_i} \pi \nabla \cdot \left(\frac{1}{\kappa} \nabla \epsilon \right) = 0 \quad (114b)$$

$$\int_{\Omega_i} \mu \left\{ \nabla \cdot \left(\frac{1}{\kappa} \nabla w \right) - \Delta_{\perp} c \right\} = 0 \quad (114c)$$

for arbitrary $(\delta, \epsilon) \in \mathcal{H}$ and arbitrary Ω_i -periodic μ satisfying (113). Integration by parts in (114a) and (114b) gives

$$\int_{\Omega_i} \left\{ -2\Delta c + R w \frac{\partial C_0}{\partial z} - \Delta_{\perp} \pi \right\} \delta = 0,$$

and

$$\int_{\Omega_i} \left\{ R c \frac{\partial C_0}{\partial z} + \nabla \cdot \left(\frac{1}{\kappa} \nabla \pi \right) \right\} \epsilon + \int_{\{z=0\}} \frac{1}{\kappa} \frac{\partial \pi}{\partial z} \epsilon = 0.$$

This implies the Euler-Lagrange equations

$$\begin{cases} -2\Delta c - R w \frac{\partial C_0}{\partial z} = \Delta_{\perp} \pi \end{cases} \quad (115a)$$

$$\begin{cases} \nabla \cdot \left(\frac{1}{\kappa} \nabla \pi \right) = -R c \frac{\partial C_0}{\partial z} \end{cases} \quad (115b)$$

$$\begin{cases} \nabla \cdot \left(\frac{1}{\kappa} \nabla w \right) = \Delta_{\perp} c \end{cases} \quad (115c)$$

in Ω_i , and

$$\frac{\partial \pi}{\partial z} \Big|_{z=0} = 0 \quad (115d)$$

as natural boundary condition.

Remark 8 Let (c, w, π, R) be a stationary point of functional (65), i.e. a nontrivial solution of (115) subject to the appropriate boundary conditions. Multiplying (115a) by c and integrating the result gives

$$2 \int_{\Omega_i} |\nabla c|^2 + R \int_{\Omega_i} wc \frac{\partial C_0}{\partial z} = \int_{\Omega_i} c \Delta_{\perp} \pi = \int_{\Omega_i} \pi \Delta_{\perp} c . \quad (116)$$

Next multiply (115c) by π . Integration yields

$$\int_{\Omega_i} \pi \nabla \left(\frac{1}{\kappa} \nabla w \right) = \int_{\Omega_i} w \nabla \cdot \left(\frac{1}{\kappa} \nabla \pi \right) = \int_{\Omega_i} \pi \Delta_{\perp} c . \quad (117)$$

In this expression we use (115b) to find

$$\int_{\Omega_i} \pi \Delta_{\perp} c = -R \int_{\Omega_i} cw \frac{\partial C_0}{\partial z} .$$

Substitution into (116) gives the identity

$$\int_{\Omega_i} |\nabla c|^2 + R \int_{\Omega_i} cw \frac{\partial C_0}{\partial z} = 0 . \quad (118)$$

We conclude that for each stationary point $\{c, w, \pi; R\}$ of (65), we have

$$J(c, w, \pi; R) = 0 .$$

E. A LOWER BOUND FOR THE ENERGY METHOD

Multiplying the equations in (P_E^*) by c , π , and w , respectively, and integrating the results gives

$$\|c'\|^2 + b^2 \|c\|^2 = \frac{R^*}{2} \int_0^\infty wce^{-\eta^2/4} - \frac{b^2}{2} \int_0^\infty \pi c , \quad (119a)$$

$$\|\pi'\|^2 + b^2 \|\pi\|^2 = -R^* \int_0^\infty c\pi e^{-\eta^2/4} , \quad (119b)$$

$$\|w'\|^2 + b^2 \|w\|^2 = b^2 \int_0^\infty cw , \quad (119c)$$

where ' denotes differentiation with respect to η . Thus every non-trivial solution of (P_E^*) satisfies

$$R^* \int_0^\infty c\pi e^{-\eta^2/4} < 0 \quad \text{and} \quad \int_0^\infty cw > 0 . \quad (120)$$

Next we multiply $(P_E^*)_2$ by w and $(P_E^*)_3$ by π . Integration gives

$$\int_0^\infty w' \pi' + b^2 \int_0^\infty w\pi = -R^* \int_0^\infty wce^{-\eta^2/4} \quad (121)$$

and

$$\int_0^\infty \pi' w' + b^2 \int_0^\infty \pi w = b^2 \int_0^\infty \pi c . \quad (122)$$

This gives the identity

$$R^* \int_0^\infty wce^{-\eta^2/4} + b^2 \int_0^\infty \pi c = 0 . \quad (123)$$

Using (123) in (119a) gives

$$\|c'\|^2 + b^2\|c\|^2 = R^* \int_0^\infty wce^{-\eta^2/4} = -b^2 \int_0^\infty \pi c . \quad (124)$$

Thus in addition to (120), we also have

$$R^* \int_0^\infty wce^{-\eta^2/4} > 0 \quad \text{and} \quad \int_0^\infty \pi c < 0 . \quad (125)$$

From (119c) we deduce

$$\|w\|^2 < \int_0^\infty cw \leq \|c\| \|w\| ,$$

or

$$\|w\| \leq \|c\| . \quad (126)$$

From (124) we deduce

$$\|c\|^2 < - \int_0^\infty \pi c \leq \|\pi\| \|c\| .$$

Thus any solution of (P_E^*) satisfies

$$\|w\| < \|c\| < \|\pi\| . \quad (127)$$

Then (119b) and (127) imply

$$R^* = 0 \quad \Rightarrow \quad \pi = c = w = 0 .$$

i.e. R^* is not in the spectrum.

Remark 9 If the eigenfunction $c > 0$ in \mathbb{R}^+ , one has (from $(P_E^*)_3$) $w > 0$ in \mathbb{R}^+ and (from $(P_E^*)_2$) $\pi < 0$ in \mathbb{R}^+ . Then (120) and (125) implies $R^* > 0$.

Identity (124) combined with inequality (126) gives

$$\begin{aligned} b^2\|c\|^2 &\leq R^* \int_0^\infty wce^{-\eta^2/4} \\ &\leq |R^*| \int_0^\infty |wc| \leq |R^*| \|w\| \|c\| \\ &< |R^*| \|c\|^2 . \end{aligned}$$

Thus $|R^*| > b^2$, implying

$$|R_E(a, t)| > \sqrt{\frac{\pi}{t}} a^2 t = a^2 \sqrt{\pi t} . \quad (128)$$

Assuming $R_E > 0$, inequality (128) yields stability for any wave number $a > 0$ and t sufficiently large.

F. SOLUTION BY GREEN'S KERNELS

We assume $\kappa(z)=1$. Let $\mathcal{G}_1(\eta;s)$ be the solution of

$$\begin{cases} -\frac{d^2\mathcal{G}_1}{d\eta} + b^2\mathcal{G}_1 = \delta(\eta-s), & \eta > 0, s > 0, \\ \frac{d\mathcal{G}_1}{d\eta}(0) = 0, & \mathcal{G}_1 \rightarrow 0 \text{ as } \eta \rightarrow \infty, \end{cases}$$

and let $\mathcal{G}_2(\eta;s)$ be the solution of

$$\begin{cases} -\frac{d^2\mathcal{G}_2}{d\eta^2} + b^2\mathcal{G}_2 = \delta(\eta-s), & \eta > 0, s > 0, \\ \mathcal{G}_2(0) = 0, & \mathcal{G}_2 \rightarrow 0 \text{ as } \eta \rightarrow \infty, \end{cases}$$

Then \mathcal{G}_1 and $\mathcal{G}_2(\eta;s)$ are given by respectively

$$\mathcal{G}_1(\eta;s) = \begin{cases} \frac{-1}{2b} (e^{b(\eta-s)} + e^{-b(\eta+s)}) & \text{for } \eta < s, \\ \frac{-1}{2b} (e^{b(s-\eta)} + e^{-b(s+\eta)}) & \text{for } \eta > s. \end{cases}$$

and

$$\mathcal{G}_2(\eta;s) = \begin{cases} \frac{-1}{2b} (e^{b(\eta-s)} - e^{-b(\eta+s)}) & \text{for } \eta < s, \\ \frac{-1}{2b} (e^{b(s-\eta)} - e^{-b(s+\eta)}) & \text{for } \eta > s. \end{cases}$$

Using these Green's functions we have from $(P_L^*)_2$ and $(P_E^*)_2$

$$w(\eta) = b^2 \int_0^\infty \mathcal{G}_1(\eta;s) c(s) ds,$$

and from $(P_E^*)_3$

$$\pi(\eta) = -R^* \int_0^\infty \mathcal{G}_1(\eta;s) e^{-s^2/4} c(s) ds.$$

Then (P_L^*) can be rewritten as

$$\begin{aligned} c(\eta) &= R^* b^2 \int_0^\infty \mathcal{G}_2(\eta;u) \left\{ \int_0^\infty \mathcal{G}_1(u;s) c(s) ds \right\} e^{-u^2/4} du \\ &= R^* b^2 \int_0^\infty \left\{ \int_0^\infty \mathcal{G}_2(\eta;u) \mathcal{G}_1(u,s) e^{-u^2/4} du \right\} c(s) ds, \end{aligned}$$

or

$$c(\eta) = \lambda_L \int_0^\infty K_L(\eta,s) c(s) ds, \tag{129}$$

with $\lambda_L = R^* b$ and $K_L(\eta;s) = \int_0^\infty \mathcal{G}_2(\eta;u) \mathcal{G}_1(u,s) e^{-u^2/4} du$.

Similarly, (P_E^*) can be written as

$$c(\eta) = \lambda_E \int_0^\infty K_E(\eta,s) c(s) ds, \tag{130}$$

where $\lambda_E = \frac{R^* b^2}{2}$ and $K_E(\eta; s) = \int_0^\infty \mathcal{G}_2(\eta; u) \mathcal{G}_1(u; s) \{e^{-u^2/4} + e^{-s^2/4}\} du$.

In an abstract way, equations (129) and (130) are written as

$$c = \lambda_L \mathcal{K}_L c \quad \text{and} \quad c = \lambda_E \mathcal{K}_E c, \quad (131)$$

where \mathcal{K}_L and \mathcal{K}_E denote the integral operators defined in (129) and (130).

Let $Q = \{(\eta, t) : \eta > 0, t > 0\}$ and $\mathbb{R}^+ = (0, \infty)$. We have the following properties of the integral operators:

- (P1) $K_i > 0$ in Q for $i = L, E$. This is a direct consequence of $\mathcal{G}_1, \mathcal{G}_2 > 0$.
- (P2) K_i are non-symmetric, i.e. $K_i(\eta, t) \neq K_i(t, \eta)$ for $i = L, E$. Although the Green's functions \mathcal{G}_1 and \mathcal{G}_2 themselves are symmetric, they are different due to the different boundary conditions: $c(0) = 0$ and $w'(0) = \pi'(0) = 0$. Consequently, $\mathcal{G}_2(\eta; s) \mathcal{G}_1(s; t) \neq \mathcal{G}_2(t; s) \mathcal{G}_1(s; \eta)$. If w and π would satisfy the Dirichlet conditions $w(0) = \pi(0) = 0$, then $\mathcal{G}_1 = \mathcal{G}_2$ and the kernels would be symmetric.
- (P3) $\|K_i\|_{L^p(Q)} < \infty$ for $i = L, E$ and for all $1 \leq p \leq \infty$. See Appendix for proof. In particular, $\|K_i\|_{L^2(Q)} < \infty$ for $i = L, E$. This means (see Renardy and Rogers (2004)) that both kernels are of Hilbert–Schmidt type and that K_L and K_E are compact operators from $L^2(\mathbb{R}^+)$ to $L^2(\mathbb{R}^+)$.
- (P4) $K_E(\eta, t) = K_L(\eta, t) + \Delta(\eta, t)$, where $\Delta(\eta, t) = e^{-t^2/4} \int_0^\infty \mathcal{G}_2(\eta, s) \mathcal{G}_1(s, t) ds$. Since $\Delta > 0$ in Q , we have $K_E > K_L$ in Q . Together with (P1) this implies $\|K_E\|_{L^2(Q)} > \|K_L\|_{L^2(Q)}$.
- (P5) Combining (131) and (P3) we have $\|c\|_{L^2(\mathbb{R}^+)}^2 = \lambda_i (K_i c, c)_{L^2(\mathbb{R}^+)}$ for $i = L, E$. This means that

$$\frac{1}{\lambda_{i,min}} = \sup_{\|c\|_{L^2(\mathbb{R}^+)}=1} (K_i c, c)_{L^2(\mathbb{R}^+)}.$$

Since $(K_i c, c)_{L^2(\mathbb{R}^+)} = \int_Q K_i(\eta, t) c(t) c(\eta) dt d\eta$, we estimate with Schwartz

$$(K_i c, c)_{L^2(\mathbb{R}^+)} \leq \left(\int_Q K_i^2 \right)^{1/2} \left(\int_Q c^2(t) c^2(\eta) d\eta dt \right)^{1/2},$$

or

$$(K_i c, c)_{L^2(\mathbb{R}^+)} \leq \|K_i\|_{L^2(Q)} \|c\|_{L^2(\mathbb{R}^+)}^2.$$

Thus

$$\frac{1}{\lambda_{i,min}} \leq \|K_i\|_{L^2(Q)}.$$

G. L^p -BOUNDEDNESS OF K_i FOR $i = L, E$

We will show that $K_i \in L^p(Q)$ for all $1 \leq p \leq \infty$ (with $i = E, L$). Starting point is the expression

$$K_E(\eta, t) = \int_0^\infty \mathcal{G}_2(\eta; s) \mathcal{G}_1(s; t) \{e^{-s^2/4} + e^{-t^2/4}\} ds. \quad (132)$$

Using the explicit expressions for \mathcal{G}_1 and \mathcal{G}_2 we construct bounds for (note that $K_E > 0$ in Q)

$$\|K_E\|_\infty = \sup_{(\eta, t) \in Q} K_E(\eta, t), \quad (133)$$

and

$$\|K_E\|_1 = \int_Q K_E(\eta, t) d\eta dt. \quad (134)$$

These bounds depend on the parameter b . From (133) and (134) it follows that

$$\|K_E\|_p = \left(\int_Q K_E^p \right)^{1/p} = \left(\int_Q K_E^{p-1} K_E \right)^{1/p} \leq \|K_E\|_\infty^{(p-1)/p} \|K_E\|_1^{1/p}$$

for $1 < p < \infty$.

Since $\mathcal{G}_2(\eta; s)$ and $\mathcal{G}_1(s; t)$ change with $\eta \leq s$ and $s \leq t$, we need to do some bookkeeping. We consider separately

$$K_E(\eta; t) = \begin{cases} K_E^+(\eta, t) & \text{in } Q^+ := \{t > \eta\} \\ K_E^-(\eta, t) & \text{in } Q^- := \{t < \eta\} \end{cases}$$

In the integral (132) we consider three intervals

$$\begin{aligned} 0 < s < \eta, \quad \eta < s < t, \quad \text{and} \quad s > t & \quad \text{when } (\eta, t) \in Q^+, \\ 0 < s < t, \quad t < s < \eta, \quad \text{and} \quad s > \eta & \quad \text{when } (\eta, t) \in Q^-. \end{aligned}$$

In Q^+ this gives

$$\begin{aligned} K_E(\eta, t) &= K_E^+(\eta, t) = \int_0^\eta \cdots + \int_\eta^t \cdots + \int_t^\infty \cdots \\ &=: I_1^+(\eta, t) + I_2^+(\eta, t) + I_3^+(\eta, t). \end{aligned} \quad (135)$$

Likewise, in Q^-

$$\begin{aligned} K_E(\eta, t) &= K_E^-(\eta, t) = \int_0^t \cdots + \int_t^\eta \cdots + \int_\eta^\infty \cdots \\ &=: I_1^-(\eta, t) + I_2^-(\eta, t) + I_3^-(\eta, t). \end{aligned} \quad (136)$$

We construct $\|\cdot\|_\infty$ and $\|\cdot\|_1$ bounds for I_i^\pm in Q^\pm ($i=1, 2, 3$).

In the estimates below we use (trivially)

$$\left. \begin{aligned} e^{-t^2/4} &< e^{-s^2/4} & \text{when } t > s \\ e^{-s^2/4} &< e^{-t^2/4} & \text{when } s > t \end{aligned} \right\} \quad (137)$$

We first estimate in Q^+ (i.e. $0 < \eta < t < \infty$). Then the product of the Green's functions can be written as

$$\mathcal{G}_2(\eta; s) \mathcal{G}_1(s; t) = \frac{1}{4b^2} \begin{cases} e^{-b(\eta+t)} (e^{2bs} - e^{-2bs}) & s < \eta < t \end{cases} \quad (138a)$$

$$\begin{cases} (e^{b(\eta-t)} - e^{-b(\eta+t)}) (1 + e^{-2bs}) & \eta < s < t \end{cases} \quad (138b)$$

$$\begin{cases} (e^{b\eta} - e^{-b\eta}) (e^{bt} + e^{-bt}) e^{-2bs} & \eta < t < s \end{cases} \quad (138c)$$

Estimates for I_1^+ : Using (138a) in (132) we estimate

$$\int_0^\eta (e^{2bs} - e^{-2bs}) (e^{-s^2/4} + e^{-t^2/4}) ds < 2 \int_0^\eta e^{2bs-s^2/4} ds = 2e^{4b^2} \int_0^\eta e^{-(s-4b)^2/4} ds < 4\sqrt{\pi} e^{4b^2}.$$

Hence

$$I_1^+(\eta, t) < \frac{\sqrt{\pi}}{b^2} e^{4b^2} e^{-b(\eta+t)},$$

implying

$$\|I_1^+\|_{L^\infty(Q^+)} \leq \frac{\sqrt{\pi}}{b^2} e^{4b^2} \quad \text{and} \quad \|I_1^+\|_{L^1(Q^+)} \leq \frac{\sqrt{\pi}}{2b^4} e^{4b^2}. \quad (139)$$

Estimates for I_2^+ : Using (138b) in (132) we estimate

$$\int_{\eta}^t (1 + e^{-2bs}) (e^{-s^2/4} + e^{-t^2/4}) ds < 4 \int_{\eta}^t e^{-s^2/4} ds .$$

Hence

$$I_2^+(\eta, t) < \frac{1}{b^2} e^{b(\eta-t)} \int_{\eta}^t e^{-s^2/4} ds ,$$

implying

$$\|I_2^+\|_{L^\infty(Q^+)} < \frac{\sqrt{\pi}}{b^2} , \quad (140)$$

and, using

$$\begin{aligned} \int_0^t I_2^+(\eta, t) d\eta &< \frac{1}{b^3} e^{b(\eta-t)} \int_{\eta}^t e^{-s^2/4} ds \Big|_0^t + \frac{1}{b^3} \int_0^t e^{b(\eta-t)-\eta^2/4} d\eta \\ &= \frac{1}{b^3} e^{-bt} \left\{ \int_0^t e^{b\eta-\eta^2/4} d\eta - \int_0^t e^{-s^2/4} ds \right\} \\ &< \frac{1}{b^3} e^{-bt} e^{b^2} \int_0^t e^{-(\eta-2b)^2/4} d\eta \\ &< \frac{2\sqrt{\pi}}{b^3} e^{b^2} e^{-bt} , \end{aligned}$$

$$\|I_2^+\|_{L^1(Q^+)} < \frac{2\sqrt{\pi}}{b^4} e^{b^2} . \quad (141)$$

Estimates for I_3^+ : Using (138c) in (132) we estimate

$$\int_t^\infty e^{-2bs} (e^{-s^2/4} + e^{-t^2/4}) ds < 2e^{-t^2/4} \int_t^\infty e^{-2bs} ds = \frac{1}{b} e^{-t^2/4-2bt} .$$

Hence

$$I_3^+(\eta, t) < \frac{1}{4b^3} e^{-t^2/4} (e^{-bt} + e^{-3bt}) (e^{b\eta} - e^{-b\eta}) ,$$

implying (since $\eta < t$)

$$\|I_3^+\|_{L^\infty(Q^+)} \leq \frac{1}{2b^3} , \quad (142)$$

and, using

$$\begin{aligned} (e^{-bt} + e^{-3bt}) \int_0^t (e^{b\eta} - e^{-b\eta}) d\eta &= \frac{1}{b} (e^{-bt} + e^{-3bt}) (e^{bt} + e^{-bt} - 2) \\ &\quad \frac{1}{b} (1 + e^{-2bt}) (1 - e^{-bt})^2 < \frac{2}{b} , \end{aligned}$$

$$\|I_3^+\|_{L^1(Q^+)} \leq \frac{\sqrt{\pi}}{2b^4} . \quad (143)$$

Next we estimate in Q^- ($0 < t < \eta < \infty$). Then we have

$$\mathcal{G}_2(\eta; s) \mathcal{G}_1(s; t) = \frac{1}{4b^2} \begin{cases} e^{-b(\eta+t)} (e^{2bs} - e^{-2bs}) & s < t < \eta \end{cases} \quad (144a)$$

$$\begin{cases} (e^{b(t-\eta)} - e^{-b(\eta+t)}) (1 - e^{-2bs}) & t < s < \eta \end{cases} \quad (144b)$$

$$\begin{cases} (e^{b\eta} - e^{-b\eta}) (e^{bt} + e^{-bt}) e^{-2bs} & t < \eta < \infty \end{cases} \quad (144c)$$

Estimates for I_1^- : Using (144a) in (132) we estimate

$$\int_0^t \left(e^{2bs} - e^{-2bs} \right) \left(e^{-s^2/4} + e^{-t^2/4} \right) ds < 4\sqrt{\pi} e^{4b^2}.$$

Hence

$$I_1^-(\eta, t) < \frac{\sqrt{\pi}}{b^2} e^{4b^2} e^{-b(\eta+t)}.$$

Consequently

$$\|I_1^-\|_{L^\infty(Q^-)} < \frac{\sqrt{\pi}}{b^2} e^{4b^2} \quad \text{and} \quad \|I_1^-\|_{L^1(Q^-)} < \frac{\sqrt{\pi}}{2b^4} e^{4b^2} \quad (145)$$

Estimates for I_2^- : Using (144b) in (132) we estimate

$$\int_t^\eta \left(1 - e^{-2bs} \right) \left(e^{-s^2/4} + e^{-t^2/4} \right) ds < 2(\eta - t) e^{-t^2/4}.$$

Thus

$$I_2^-(\eta, t) < \frac{1}{2b^2} \left(e^{b(t-\eta)} - e^{-b(t+\eta)} \right) (\eta - t) e^{-t^2/4},$$

implying

$$\|I_2^-\|_{L^\infty(Q^-)} < \frac{1}{2eb^3}, \quad (146)$$

and, using

$$\begin{aligned} \int_t^\infty I_2^-(\eta, t) d\eta &< \frac{1}{2b^2} \int_t^\infty (\eta - t) e^{-b(\eta-t)-t^2/4} d\eta \\ &= \frac{1}{2b^2} \int_0^\infty \zeta e^{-b\zeta} d\zeta e^{-t^2/4} = \frac{1}{4b^4} e^{-t^2/4}, \end{aligned}$$

$$\|I_2^-\|_{L^1(Q^-)} < \frac{\sqrt{\pi}}{2b^4}. \quad (147)$$

Estimates for I_3^- : Using (144c) in (132) we estimate

$$\int_\eta^\infty e^{-2bs} \left(e^{-s^2/4} + e^{-t^2/4} \right) ds < 2e^{-t^2/4} \int_\eta^\infty e^{-2bs} ds = \frac{1}{b} e^{-t^2/4-2b\eta}.$$

Hence

$$\begin{aligned} I_3^-(\eta, t) &< \frac{1}{4b^3} \left(e^{b\eta} - e^{-b\eta} \right) \left(e^{bt} + e^{-bt} \right) e^{-t^2/4-2b\eta} \\ &= \frac{e^{b(\eta+t)}}{4b^3} \left(1 - e^{-2b\eta} \right) \left(1 + e^{-2bt} \right) e^{-t^2/4-2b\eta} < \frac{1}{2b^3} e^{bt-t^2/4-b\eta}. \end{aligned}$$

Consequently

$$\|I_3^-\|_{L^\infty(Q^-)} < \frac{1}{2b^3} \quad \text{and} \quad \|I_3^-\|_{L^1(Q^-)} < \frac{\sqrt{\pi}}{b^4} e^{b^2}. \quad (148)$$

Finally we combine (135)–(148) to find

$$\begin{aligned}\|K_E^+\|_{L^\infty(Q^+)} &\leq \frac{\sqrt{\pi}}{b^2} e^{4b^2} + \frac{\sqrt{\pi}}{b^2} + \frac{1}{2b^3} , \\ \|K_E^-\|_{L^\infty(Q^-)} &\leq \frac{\sqrt{\pi}}{b^2} e^{4b^2} + \frac{1}{2eb^3} + \frac{1}{2b^3} ,\end{aligned}$$

implying

$$\|K_E\|_{L^\infty(Q)} \leq \frac{\sqrt{\pi}}{b^2} e^{4b^2} + \frac{1}{2b^3} + \frac{\max\{\sqrt{\pi}, \frac{1}{2eb}\}}{b^2} , \quad (149)$$

and

$$\|K_E\|_{L^1(Q)} \leq \frac{\sqrt{\pi}}{b^4} e^{4b^2} + \frac{3\sqrt{\pi}}{b^4} e^{b^2} + \frac{\sqrt{\pi}}{b^4} . \quad (150)$$

This means that

$$\|K_E\|_{L^2(Q)} \leq \sqrt{\|K_E\|_{L^\infty(Q)} \cdot \|K_E\|_{L^1(Q)}} \leq F(b) ,$$

where $F(b)$ is the square root of the product of the right sides in (149) and (150).

Remark 10 Replacing in K_E the exponential functions $e^{-s^2/4}$ and $e^{-t^2/4}$ by a constant, yields an unbounded L^1 (or L^2) norm. Then the corresponding integral operator is unbounded.

Remark 11 In all estimates we used (137). As a consequence the same result holds for the kernel $K_L(\eta, t)$. The only difference with the estimates above is a factor $\frac{1}{2}$. This is then compensated by the difference in the definitions of λ_E and λ_L .

H. CHEBYSHEV–PETROV–GALERKIN SPECTRAL METHOD

H.1. Linearised perturbation equations

The linearised perturbation equations (43) are solved by means of a Chebyshev–Petrov–Galerkin spectral method, see Shen et al. (2011). Spectral methods deal with weighted residual problems. We first introduce two linear operators:

$$L' := \alpha \frac{\partial^2}{\partial \zeta^2} - a^2 , \quad L := -\alpha \frac{\partial}{\partial \zeta} \left(\frac{1}{\widehat{\kappa}(\zeta)} \frac{\partial}{\partial \zeta} \right) + \frac{a^2}{\widehat{\kappa}(\zeta)} .$$

The $\widehat{\cdot}$ -notation and parameter $\alpha = 4/d^2$ stem from the trivial linear coordinate transformation which maps the interval $(0, d)$ to $I := (-1, 1)$.

Starting point of the weighted residual method is to approximate the solution c and w of (43) by a finite sum

$$\begin{aligned}c(z, t) &= \widehat{c}(\zeta, t) \approx \widehat{c}_N(\zeta, t) = \sum_{i=0}^N c_i(t) \phi_i(\zeta) , \\ w(z, t) &= \widehat{w}(\zeta, t) \approx \widehat{w}_N(\zeta, t) = \sum_{i=0}^N w_i(t) \varphi_i(\zeta) ,\end{aligned}$$

where $\{\phi_i\}$ and $\{\varphi_i\}$ are the trial (or shape) functions. The coefficients $\{c_i(t)\}$ and $\{w_i(t)\}$ are to be determined by solving the following problem:

$$(CPG-I) \left\{ \begin{array}{l} \text{Find } \{c_i(t)\} \text{ and } \{w_i(t)\} \text{ for } t > 0 \text{ such that:} \\ \int_I \left\{ \left[L' - \frac{\partial}{\partial t} \right] \sum_{i=0}^N c_i(t) \phi_i - R \sqrt{\alpha} \frac{\partial \widehat{C}_0}{\partial \zeta} \sum_{i=0}^N w_i(t) \varphi_i \right\} \psi_j \omega = 0, \quad 0 \leq j \leq N, \\ \int_I \left\{ L \sum_{i=0}^N w_i(t) \varphi_i - a^2 \sum_{i=0}^N c_i(t) \phi_i \right\} \psi_j \omega = 0, \quad 0 \leq j \leq N, \\ \int_I \left\{ \sum_{i=0}^N c_i(0) \phi_i - \widehat{c}(\cdot, 0) \right\} \psi_j \omega = 0, \quad 0 \leq j \leq N. \end{array} \right.$$

Here $\{\psi_j\}$ are the test functions, and ω a positive weight function.

In our approach we define the shape functions by (Heinrichs, 1989, 1991; Shen et al., 2011)

$$\begin{aligned} \phi_i(\zeta) &= (1 - \zeta^2) T_i(\zeta), \\ \varphi_i(\zeta) &= T_i(\zeta) + a_i T_{i+1}(\zeta) + b_i T_{i+2}(\zeta), \end{aligned}$$

and the test functions by (Gottlieb and Orszag, 1977; Shen, 1995)

$$\psi_j(\zeta) = T_j(\zeta) - T_{j+2}(\zeta).$$

The coefficients $\{a_i\}$ and $\{b_i\}$ are chosen in such a way that the shape function for \widehat{w}_N satisfies the Neumann boundary condition $\frac{\partial \widehat{w}_N}{\partial z}(-1, \cdot) = 0$ ($z=0$) and the Dirichlet condition $\widehat{w}_N(1, \cdot) = 0$ ($z=H$):

$$a_i = \frac{4i + 4}{-2i^2 - 6i - 5}, \quad \text{and} \quad b_i = \frac{2i^2 + 2i + 1}{-2i^2 - 6i - 5}.$$

The shape function for \widehat{c}_N satisfies the Dirichlet conditions $\widehat{c}_N(-1, \cdot) = \widehat{c}_N(1, \cdot) = 0$.

Since both the shape and test functions are build up from Chebyshev polynomials, we will use for the weighting function ω

$$\omega(\zeta) = \frac{1}{\sqrt{1 - \zeta^2}}$$

as to exploit as much as possible the orthogonality property.

We define the following matrices:

$$\begin{aligned} \mathbf{A}_{i,j} &:= \alpha \int_I \frac{d^2 \phi_i}{d\zeta^2} \psi_j \omega, \quad \mathbf{I}_{i,j} := \int_I \phi_i \psi_j \omega, \quad \mathbf{J}_{i,j} := \int_I \frac{1}{\widehat{\kappa}(\zeta)} \varphi_i \psi_j \omega, \\ \mathbf{K}_{i,j} &:= \alpha \int_I \frac{d}{d\zeta} \left(\frac{1}{\widehat{\kappa}(\zeta)} \frac{d\varphi_i}{d\zeta} \right) \psi_j \omega = -\alpha \int_I \frac{1}{\widehat{\kappa}(\zeta)} \frac{d\varphi_i}{d\zeta} \frac{d\psi_j}{d\zeta} \omega - \alpha \int_I \frac{1}{\widehat{\kappa}(\zeta)} \frac{\zeta}{1 - \zeta^2} \frac{d\varphi_i}{d\zeta} \psi_j \omega, \\ \mathbf{P}_{i,j}(t) &:= \sqrt{\alpha} \int_I \frac{\partial \widehat{C}_0}{\partial \zeta}(\cdot, t) \varphi_i \psi_j \omega, \\ \mathbf{f}_j &:= \int_I \widehat{c}(\cdot, 0) \psi_j \omega. \end{aligned}$$

In this formulation we allow for non-differentiable, but non-zero $\widehat{\kappa}$.

The Chebyshev–Gauss quadrature (Quarteroni et al., 2000) is an extension of Gaussian quadrature method for approximating the value of integrals of the following kind:

$$\int_I f \omega = \int_I \frac{f(\zeta)}{\sqrt{1-\zeta^2}} d\zeta .$$

Its value can be approximated by

$$\int_I \frac{f(\zeta)}{\sqrt{1-\zeta^2}} d\zeta \approx \frac{\pi}{M} \sum_{i=1}^M f\left(\cos\left(\frac{2i-1}{2M}\pi\right)\right) .$$

Let $\mathbf{c} = (c_0(t), c_1(t), \dots, c_N(t))^T$ and $\mathbf{w} = (w_0(t), w_1(t), \dots, w_N(t))^T$. Then problem (CPG-I) can be written in matrix-vector notation:

$$\begin{cases} \text{Find } \mathbf{c}(t) \text{ and } \mathbf{w}(t) \text{ for } t > 0 \text{ such that:} \\ \left[\mathbf{A} - a^2 \mathbf{I} \right] \mathbf{c} - \mathbf{I} \frac{d\mathbf{c}}{dt} - R\mathbf{P}(t)\mathbf{w} = \mathbf{0} , \\ \left[-\mathbf{K} + a^2 \mathbf{J} \right] \mathbf{w} - a^2 \mathbf{I} \mathbf{c} = \mathbf{0} , \\ \mathbf{I} \mathbf{c}(0) - \mathbf{f} = \mathbf{0} . \end{cases}$$

To analyse linear stability, we assume $\frac{d\mathbf{c}}{dt} = \sigma \mathbf{c}$ and solve the problem:

$$\begin{cases} \text{For a given } t > 0, \text{ find } \mathbf{c}(t) \text{ and } \mathbf{w}(t) \text{ such that:} \\ \left[\mathbf{A} - a^2 \mathbf{I} \right] \mathbf{c} - R\mathbf{P}(t)\mathbf{w} = \sigma \mathbf{I} \mathbf{c} , \\ \left[-\mathbf{K} + a^2 \mathbf{J} \right] \mathbf{w} - a^2 \mathbf{I} \mathbf{c} = \mathbf{0} . \end{cases}$$

H.2. The Euler–Lagrange equations from the energy method

By the nature of the (implied) boundary conditions for π , we approximate its solution by

$$\pi(z, t) = \widehat{\pi}(\zeta, t) \approx \widehat{\pi}_N(\zeta, t) = \sum_{i=0}^N \pi_i(t) \varphi_i(\zeta) .$$

The argument t has to be considered here as a given and fixed time parameter. Proceeding as in the previous subsection, we want to solve the following problem:

$$\text{(CPG-II)} \begin{cases} \text{For a given } t > 0, \text{ find } \{c_i(t)\}, \{w_i(t)\}, \text{ and } \{\pi_i(t)\} \text{ such that:} \\ \int_I \left\{ L' \sum_{i=0}^N c_i(t) \phi_i - \frac{1}{2} R \sqrt{\alpha} \frac{\partial \widehat{C}_0}{\partial \zeta} \sum_{i=0}^N w_i(t) \varphi_i - \frac{1}{2} a^2 \sum_{i=0}^N \pi_i(t) \varphi_i \right\} \psi_j \omega = 0 , & 0 \leq j \leq N , \\ \int_I \left\{ L \sum_{i=0}^N \pi_i(t) \varphi_i + R \sqrt{\alpha} \frac{\partial \widehat{C}_0}{\partial \zeta} \sum_{i=0}^N c_i(t) \phi_i \right\} \psi_j \omega = 0 , & 0 \leq j \leq N , \\ \int_I \left\{ L \sum_{i=0}^N w_i(t) \varphi_i + a^2 \sum_{i=0}^N c_i(t) \phi_i \right\} \psi_j \omega = 0 , & 0 \leq j \leq N , \end{cases}$$

We define the matrices

$$\mathbf{H}_{i,j} := \int_I \varphi_i \psi_j \omega \quad \text{and} \quad \mathbf{Q}_{i,j}(t) := \sqrt{\alpha} \int_I \frac{\partial \widehat{C}_0}{\partial \zeta}(\cdot, t) \phi_i \psi_j \omega .$$

Let $\boldsymbol{\pi} = (\pi_0(t), \pi_1(t), \dots, \pi_N(t))^T$. Then problem (CPG-II) can be written in matrix-vector notation:

$$\left\{ \begin{array}{l} \text{For a given } t > 0, \text{ find } c(t), w(t), \text{ and } \boldsymbol{\pi}(t) \text{ such that:} \\ \left[\mathbf{A} - a^2 \mathbf{I} \right] c - \frac{1}{2} a^2 \mathbf{H} \boldsymbol{\pi} - \frac{1}{2} R \mathbf{P}(t) w = \mathbf{0} , \\ \left[-\mathbf{K} + a^2 \mathbf{J} \right] \boldsymbol{\pi} - R \mathbf{Q}(t) c = \mathbf{0} , \\ \left[-\mathbf{K} + a^2 \mathbf{J} \right] w - a^2 \mathbf{I} c = \mathbf{0} . \end{array} \right.$$

I. CHARACTERISTIC-GALERKIN FINITE ELEMENT METHOD

The solution of (77–78) is approximated by means of a characteristic–Galerkin finite element approach (Pironneau (1982); Pironneau et al. (1992); Xikui and Wenhua (1999)). Key feature of this approach is the treatment of the convective terms, which might become dominant compared to the diffusive term for the case of non-steady convective flows. Convective dominance might give rise to numerical instabilities during time-integrations.

As with any Galerkin method, the formulation starts with a weak formulation. Let:

$$\begin{aligned} H_{\text{per}}^1 &:= \left\{ v \in H^1(\widehat{\Omega}) : v|_{\Gamma_2} = v|_{\Gamma_4}, \frac{\partial v}{\partial x}|_{\Gamma_2} = \frac{\partial v}{\partial x}|_{\Gamma_4} \right\} , \\ H_{0,\text{per}}^1 &:= \left\{ v \in H_{\text{per}}^1(\widehat{\Omega}) : v|_{\Gamma_1} = v|_{\Gamma_3} = 0 \right\} , \\ V_{0,\text{per}} &:= \left\{ v \in H_{\text{per}}^1(\widehat{\Omega}) : v|_{\Gamma_1} = 0 \right\} , \\ V_{1,\text{per}} &:= \left\{ v \in V_{0,\text{per}}(\widehat{\Omega}) : v|_{\Gamma_3} = 1 \right\} . \end{aligned}$$

We want to solve the following variational problem in domain $\widehat{\Omega}$ and for $t \in (0, T]$, $T > 0$:

$$\text{(FEP)} \left\{ \begin{array}{l} \text{Find } C \text{ in } L^2(0, T; V_{1,\text{per}}(\widehat{\Omega})), \partial_t C \text{ in } L^2(0, T; H_{1,\text{per}}^{-1}(\widehat{\Omega})) \\ \text{and } \Psi \text{ in } V_{0,\text{per}}(\widehat{\Omega}) \text{ such that} \\ \int_{\widehat{\Omega}} \frac{1}{\kappa(z)} \nabla \Psi \cdot \nabla w + \int_{\widehat{\Omega}} \frac{\partial C}{\partial x} w = 0 \quad \forall w \in H_{0,\text{per}}^1(\widehat{\Omega}) , \\ \int_{\widehat{\Omega}} \frac{\partial C}{\partial t} v + R \int_{\widehat{\Omega}} \left(-\frac{\partial \Psi}{\partial z} \frac{\partial C}{\partial x} + \frac{\partial \Psi}{\partial x} \frac{\partial C}{\partial z} \right) v + \int_{\widehat{\Omega}} \nabla C \cdot \nabla v = 0 \quad \forall v \in H_{0,\text{per}}^1(\widehat{\Omega}) . \end{array} \right.$$

Let Ψ be given for some time t , and let $\mathbf{X}(t) : (0, T] \rightarrow \mathbb{R}^2$ be the characteristic curve given by

$$\frac{d\mathbf{X}(t)}{dt} = \mathbf{u}(\mathbf{X}(t), t) , \quad \mathbf{X}(0) = \mathbf{X}_0 ,$$

and where we assume $\mathbf{u} = (-R \frac{\partial \Psi}{\partial z}, R \frac{\partial \Psi}{\partial x})$. Then

$$\begin{aligned} \frac{d}{dt} C(\mathbf{X}(t), t) &= \nabla C(\mathbf{X}(t), t) \cdot \frac{d\mathbf{X}(t)}{dt} + \frac{\partial C(\mathbf{X}(t), t)}{\partial t} \\ &= \nabla C(\mathbf{X}(t), t) \cdot \mathbf{u}(\mathbf{X}(t), t) + \frac{\partial C(\mathbf{X}(t), t)}{\partial t} \end{aligned}$$

Along the characteristic, the material derivative in the weak formulation can hence be written as:

$$\begin{aligned} \int_{\widehat{\Omega}} \frac{\partial C}{\partial t} v + R \int_{\widehat{\Omega}} \left(-\frac{\partial \Psi}{\partial z} \frac{\partial C}{\partial x} + \frac{\partial \Psi}{\partial x} \frac{\partial C}{\partial z} \right) v &= \int_{\widehat{\Omega}} \frac{dC(\mathbf{X}(t), t)}{dt} v \\ &\approx \int_{\widehat{\Omega}} \frac{S(\mathbf{X}(t), t) - C(\mathbf{X}(t - \tau), t - \tau)}{\tau} v, \\ &\approx \int_{\widehat{\Omega}} \frac{S(\mathbf{X}(t), t) - C(\mathbf{X}(t) - \mathbf{u}\tau, t - \tau)}{\tau} v. \end{aligned}$$

We define $t^n = t^{n+1} - \tau$, $t^n < t^{n+1}$, and $C^n = C(\cdot, t^n)$. Then we have

$$\int_{\widehat{\Omega}} \frac{\partial C}{\partial t} v + R \int_{\widehat{\Omega}} \left(-\frac{\partial \Psi}{\partial z} \frac{\partial C}{\partial x} + \frac{\partial \Psi}{\partial x} \frac{\partial C}{\partial z} \right) v \approx \int_{\widehat{\Omega}} \frac{C^{n+1} - \bar{C}^n}{\tau} v,$$

where \bar{C}^n denotes the interpolated solution of S^n at $\mathbf{X}(t^{n+1}) - \mathbf{u}\tau$.

We approximate the space of admissible functions v by $V_h = \{e \in H^1(\widehat{\Omega}) : \forall K \in \mathcal{T}_h, e|_K \in \mathbb{P}_2\}$, where \mathbb{P}_2 denotes the space of polynomials on \mathbb{R}^2 of degree ≤ 2 . Similarly, for w we have $W_h = \{e \in H^1(\widehat{\Omega}) : \forall K \in \mathcal{T}_h, e|_K \in \mathbb{P}_2\}$. By using the θ -scheme, we arrive at the following problem:

$$(\text{FEP})_h \left\{ \begin{array}{l} \text{Find } C_h^{n+1} \in V_h \text{ and } \Psi_h \in W_h \text{ such that for all} \\ v_h \in V_h \text{ and } w_h \in W_h \text{ we have:} \\ \int_{\widehat{\Omega}} \frac{1}{\kappa(z)} \nabla \Psi_h \cdot \nabla w_h + \int_{\widehat{\Omega}} \frac{\partial C_h^n}{\partial x} w_h = 0, \\ \int_{\widehat{\Omega}} \frac{C_h^{n+1} - \bar{C}_h^n}{\tau} v_h + \theta \int_{\widehat{\Omega}} \nabla C_h^{n+1} \cdot \nabla v_h + (1 - \theta) \int_{\widehat{\Omega}} \nabla C_h^n \cdot \nabla v_h = 0. \end{array} \right.$$

Above scheme has been implemented in `freefem++` (Hecht, 2012).

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